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Photogrammetry  
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*W.M.*

NATIONAL AERONAUTICS AND SPACE ADMINISTRATION

APOLLO 15 SIM BAY PHOTOGRAPHIC  
EQUIPMENT AND MISSION SUMMARY

MANNED SPACECRAFT CENTER  
HOUSTON, TEXAS

PREPARED BY  
MAPPING SCIENCES BRANCH  
NASA MANNED SPACECRAFT CENTER  
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## SUMMARY

The Apollo 15 mission will provide the first mapping quality photographic products of the lunar surface. This will be accomplished using a Fairchild 3-inch metric mapping camera with associated stellar camera, and ITEK optical-bar 24-inch panoramic camera and an RCA ruby laser altimeter. The purpose of this document is to present selected pertinent data concerning the photographic experiments in the Scientific Instrument Module (SIM) bay. Included is a brief description of each experiment with available calibration data, an explanation of the photographic orbital support data and the mission photographic profile of planned/accomplished objectives.

This mission summary is intended to provide overall knowledge of the experiments and a quick reference to the planned and accomplished Apollo 15 SIM bay photographic objectives. This document will be supplemented by publication of the Apollo 15 Photographic Index and 16mm microfilm of the orbital support data. These additional data will not be available for several months after the mission.

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## 1.0 INTRODUCTION

The three photographic experiments discussed in this document, the Fairchild three-inch metric camera, ITEK 24-inch optical bar panoramic camera, and the RCA ruby laser altimeter, are located in the Scientific Instrument Module (SIM) located in bay 1 of the Service Module (Fig. 1-1). The SIM bay also contains the spectrometer experiments.

During blastoff and translunar coast, the SIM bay is covered by the SIM bay door. This door is to be jettisoned approximately 4 1/2 hours before lunar orbit insertion (LOI). After door jettison, the spacecraft attitude must be controlled to prevent impingement of direct sun rays within the look-angle envelope of the optics of each sun sensitive instrument. This includes both the mapping and panoramic cameras as well as certain of the spectrometers.

Operation of the SIM bay experiments is controlled by an astronaut using three control panels 181, 278, and 230 (Figs. 1-2 and 1-2a) located on the right hand side of the spacecraft. Explanation of specific switches that affect equipment operation are explained in the text of the individual experiments. Further explanation of the control panels is not germane to the purpose of this document and would not add substantially to knowledge of the experiments.

Also, for real time monitoring of the experiments, certain telemetered data are displayed on the closed circuit TV in the Mission Operations Center, Bldg. 30, MSC (Fig. 1-3). Recommended operational limitations for the data are also given. The limitations are those valid at the printing of this summary and may be subject to change. At times the experiments may be operated outside these limitations due to mission considerations, with an

**SCIENTIFIC INSTRUMENT MODULE  
(APOLLO 15 CONFIGURATION)**

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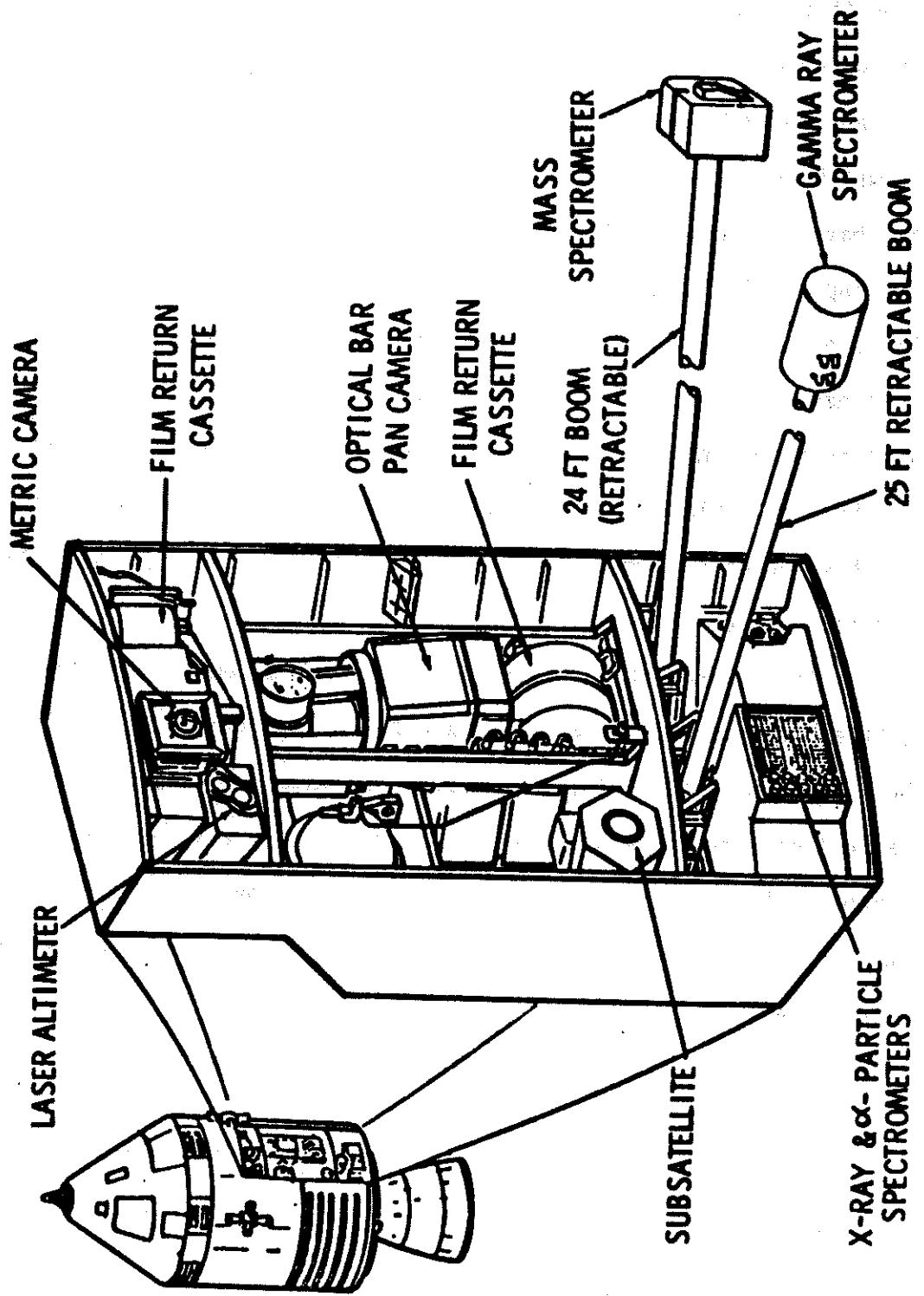
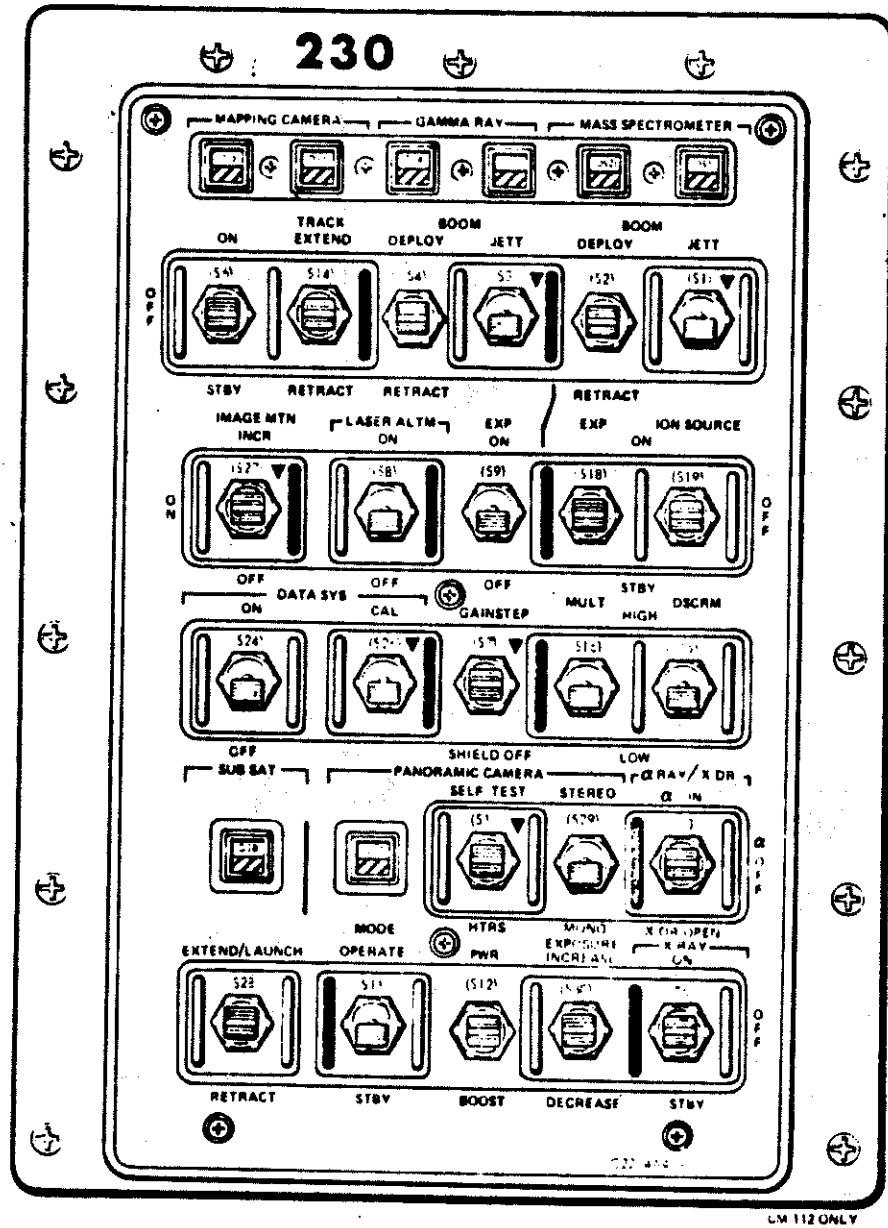


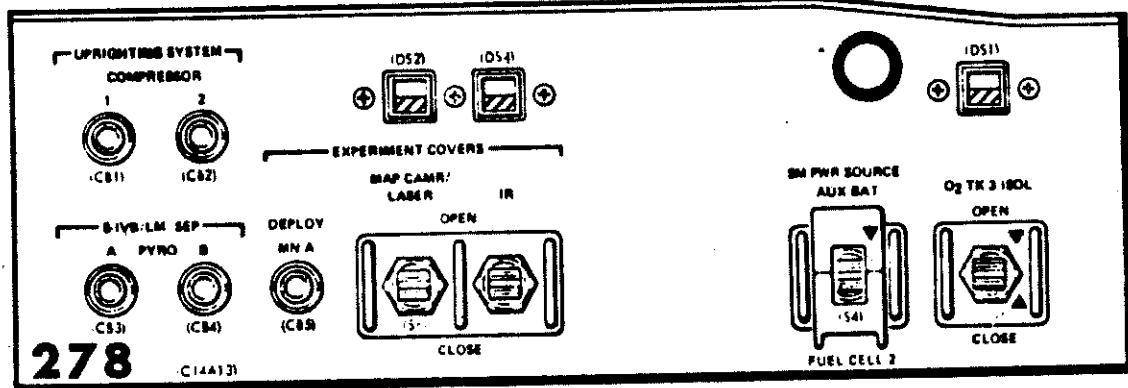
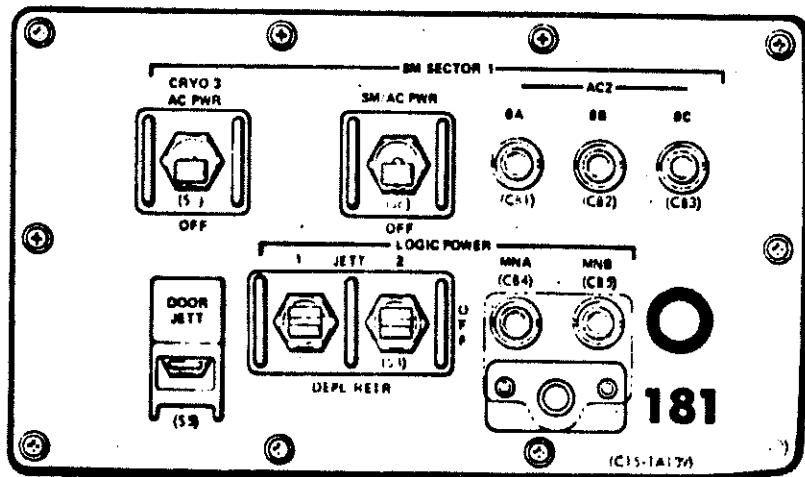
Fig. 1-1 Scientific Instrument Module Configuration

Fig. 1-1



CONTROL PANEL

Fig. 1-2



### CONTROL PANELS

Fig. 1-2a

PHOTO/LASER

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PANORAMIC CAMERA

SL1030V V/h CMD V	10.1 to 20 MRAD/Sec
SL1031X Air Sol	Event only: Open-close
SL1032T Elect Temp	86° to 110°F
SL1033M FR Roll Pos	0 to 47 inches
SL1034W Shuttle Pos	0 to 14.38 inches
SL1035C Lens Torq I	0-10 amps
SL1036X Cap Sht Pos	Event only: Open-close
SL1037V FMC Tach V	0-20 MRAD/Sec
SL1038H Expo CMD	0-2000 FL
SL1039T Lens Bar T	88° ± 2°F
SL1040T Fwd Lens T	
SL1041T Aft Lens T	
SL1042T Mech T	86° to 94°F
SL1044HS1lit-Width	.015 to .300 inches

MAPPING CAMERA

SL1160T MC Fwd Len T	
SL1161T MC Len Bar T	
SL1162T SC Fwd Len T	
SL1163T SC Len Bar T	
SL1164T Cassette T	
SL1165X Image Mot	Event only: 0 or 5V
SL1166A Shut Speed	80 to 1280 RPM
SL1166X Deploy/Cut	Event only: deploy 3.IV,
SL1176Q Film RMG	1500 to 0 ft/cut 1.0V
SL1181V V/h Inc LVL	12.1, 13.1, 14.1, 15.1, or 16.1 MRAD/Sec

LASER ALTIMETER

Reg +5V	Event only:
Ang Overflow	Condition "1" Operational
Laser Output	Condition "0" non-opera.
Q-SW Mtr Sta	
Krytron Trig	
MC Data Req	xxxxxx meters
SL1122K Altitude	Event only: 0v or 5V
SL1091V - 5V Reg	
SL1092V Photomult. V	
SL1093V PRN V	
SL1094T Cavity T	

FIG. 1-3 Photo/Laser Real Time Telemetry Display and Operational Values

accepted loss of accuracy in the resulting data.

The orientation of the experiments with respect to the lunar surface is determined from the spacecraft trajectory and onboard gimbal angles. These parameters have certain inherent errors that must be recognized. These errors, as well as others that may not be listed, affect the capability of recovering the experiment attitudes computed in the orbital support data. An explanation of attitude pointing terminology and errors is explained in Table 1-1. The orbital support data is discussed in paragraph 5.

The information in the Photographic Mission Summary has been kept brief because of the planned publication of a Photographic Data Utilization Manual for the SIM bay photographic experiments. This Manual is being compiled at the Manned Spacecraft Center and publication is planned prior to Apollo 16. If additional information related to this summary is desired please contact the Mapping Sciences Branch TF5, Manned Spacecraft Center, Houston.

## POINTING ACCURACY AND POINTING KNOWLEDGE

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### DEFINITION OF TERMS

Two major types of attitude pointing requirements are specified for each experiment's pointing accuracy and pointing knowledge.

These terms are defined as follows:

1. Pointing Accuracy Requirement/Capability - The real-time accuracy to which a given instrument must/can be pointed relative to the desired direction.

2. Pointing Knowledge Requirement/Capability - The accuracy to which a given instrument pointing history must/can be derived from post-flight data reduction.

The only significant quantitative difference between pointing accuracy capability and pointing knowledge capability is the magnitude of the attitude hold deadband maintained in lunar orbit. The instantaneous position in the control deadband can be determined from the telemetered IMU gimbal angles, thus providing a more accurate post-flight pointing knowledge.

### SPECIAL TERMS FOR BOOM-MOUNTED INSTRUMENTS

Each boom-mounted experiment presents a special pointing accuracy problem due to the additional sources of attitude uncertainty associated with the deployment mechanism and the boom itself. Hence, for the sake of convenience on these experiments, two types of pointing accuracy terms are frequently used:

1. Sensor Pointing Accuracy - Pointing accuracy related to the boom-mounted sensor, including uncertainties due to the extended boom and deployment mechanism.

## POINTING ACCURACY AND POINTING KNOWLEDGE

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2. SIM Pointing Accuracy - Pointing accuracy related to the interface of the SIM shelf and the deployment mechanism, excluding uncertainties due to the extended boom and deployment mechanism.

In that the angular uncertainties associated with the boom deployment or jettison mechanism are an order of magnitude larger than the control deadband, knowledge of the CSM attitude in the deadband is of negligible significance. Thus, for boom-mounted instruments, post-flight pointing knowledge is effectively the same as in-flight pointing accuracy.

### SOURCES OF ATTITUDE UNCERTAINTIES

The following items constitute the various sources of attitude uncertainties which contribute to the total pointing knowledge capability of all non-boom-mounted experiments:

	Attitude Uncertainty (degrees)
Alignment in the instrument*	±0.50
SIM structural tolerances	±0.50
Instrument mounting	±0.20
SIM installation tolerance	±0.23
SIM- G&N navigation base boresighting	±0.05
Deviation from one g to zero g	±1.00
Thermal-induced variations	±0.33
IMU alignment accuracy	±0.03
IMU drift (assuming six hours)	±0.54
CMC local vertical program	±0.25
Root sum of squares (RSS) of random tolerances of these 10 items	±1.44

\*Deviations to this 0.50-degree instrument design requirement may be granted for individual experiments and will be specified in the appropriate ICD's.

For non-boom-mounted experiments, the pointing accuracy capability is the 1.44 degrees plus the attitude control deadband (nominally ±0.5 degrees), or ±1.94 degrees. For the boom-mounted instruments, both pointing

TABLE 1-1 (Cont'd)

## POINTING ACCURACY AND POINTING KNOWLEDGE

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knowledge and pointing accuracy are further degraded by uncertainties in the deployment mechanism and the boom itself.

### OPERATIONAL CONSIDERATIONS

Only one planning tool is available that significantly affects pointing accuracy capability, i.e., the selection of control deadband width. For those experiments that require a total pointing accuracy of  $\pm 2$  degrees, a control deadband of  $\pm 0.5$  degrees must be selected. During those experiment periods limited to operation of instruments that require a total pointing accuracy of  $\pm 5.0$  degrees, the control deadband can be opened up to  $\pm 3.5$  degrees if desired.

TABLE 1-1 (Cont'd)

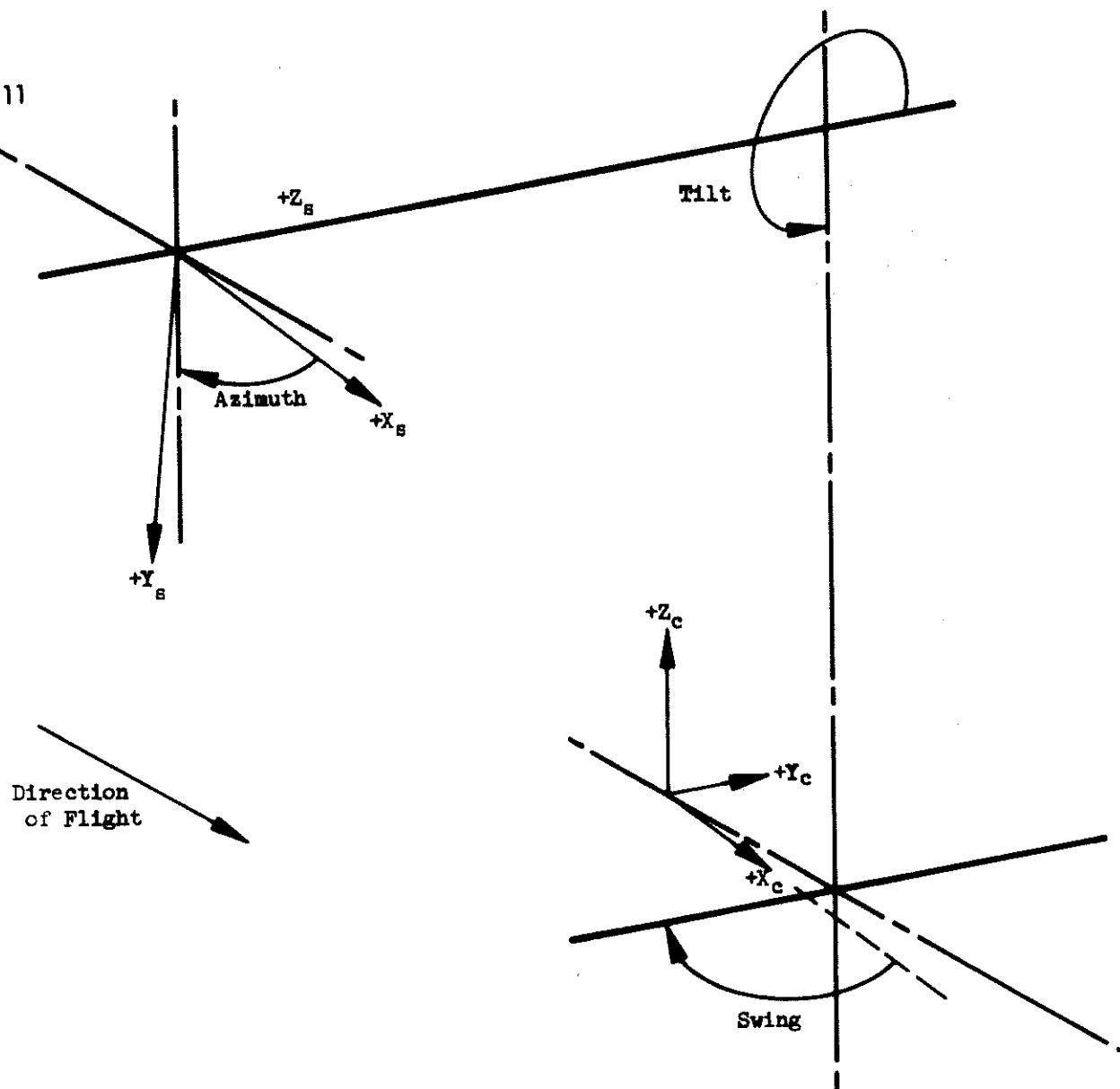
## 2.0 MAPPING CAMERA SUBSYSTEM (MCS)

The Mapping Camera Subsystem (MCS) consists of a 3-inch mapping camera utilizing 5-inch film, a 3-inch stellar camera using 35mm film, detachable supply and takeup cassettes, lens glare shields, a laser altimeter and a deployment mechanism. The interlock angle between the cameras is  $96^\circ \pm 30'$  with their angular relationship defined in terms of tilt, swing, and azimuth. (Fig. 2-1, and Table 2-1).

The MCS and its deployment mechanism are mounted on the forward side of the SIM shelf parallel to the centerline. When deployed for operational photography the parallel relationship between the mapping camera optical axis and the SIM centerline is maintained. In the event the system fails to extend, mapping photography can be accomplished, however, stellar photography cannot.

Basic camera parameters are given in Table 2-2. Additional information related to the MCS is listed below.

- The subsystem is 8 x 22 x 51 inches and weighs 280 pounds including film, record container, laser altimeter and deployment mechanism. The film record container is 11 inches in diameter by 9 inches wide and weighs 23 pounds when it contains all of the exposed film.
- The deployment mechanism is an electro-mechanical device used to extend the MCS from the SIM during lunar orbit. The stellar lens glare shield is mechanically linked to the deployment mechanism and is extended during the deployment process. Deployment, as well as retraction, requires a maximum of four minutes.



Intersection of Principal  
Plane and Carto Camera  
Focal Plane

- NOTES: (a) Carto Camera Coordinate System Rotated Slightly About  $Z_c$  Axis for Clarity  
 (b) Stellar Camera Coordinate System Rotated Slightly About  $Z_s$  Axis for Clarity

Figure 2-1 - MSG Interlock Angle Geometry

## DEFINITION OF MCS INTERLOCK ANGLE GEOMETRY

The positive X axis of the cartographic camera shall be defined as being approximately in the direction of the vehicle velocity vector. The Z axis shall be perpendicular to the X-Y plane and positive in the direction of the terrain to vehicle radius vector. The cartographic camera Y axis shall be approximately perpendicular to the orbit plane and positive in such a direction as to conform to a right hand Cartesian Coordinate System. The positive X axis of the Stellar Camera shall be approximately in the direction of the vehicle velocity vector. The negative Y axis shall be in the approximate direction of the terrain to vehicle radius vector. The Z axis of the Stellar Camera shall be perpendicular to the X-Y plane and positive in such a direction as to conform to a right hand Cartesian Coordinate System.

Within each focal plane, the X and Y axes shall be defined as follows:

The Y axis shall be defined as the line determined by the two fiducial marks which are approximately perpendicular to the vehicle velocity vector. The X Axis shall be defined as a line which is perpendicular to the Y axis, but passes through the indicated principal point. The Z axis is then defined as being perpendicular to the intersection of X and Y.

The principal plane shall be defined as that plane which contains  $Z_S$  and is parallel to  $Z_c$ .

The three angles defining the system geometry are defined as follows:

TILT	The clockwise angle measured from $Z_S$ to the projection of $Z_c$ on the principal plane (looking in the direction of flight). This angle shall be $264^\circ \pm 30$ minutes.
SWING	The angle measured from $+X_c$ to the trace of the principal plane in the -Y direction on the cartographic camera focal plane. This angle shall be $90^\circ \pm 30$ minutes.
AZIMUTH	The angle measured in the clockwise direction from $+X_S$ to the trace of the principal plane (looking in the $-Z_S$ direction). This angle shall be $90^\circ \pm 30$ minutes.

TABLE 2-1 MSC Interlock Geometry

## MAPPING CAMERA CHARACTERISTICS

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Type	Stationary film, mapping
Mode of Operation	Autocycle
Lens and Aperture	
Stellar	3 in. f/2.8 (fixed)
Mapping	3 in. f/4.5 (fixed)
Format	
Stellar	1.25 in. diameter with 0.96 in. flats
Mapping	4-1/2 x 4-1/2 in.
Coverage (Mapping)	74° x 74°
Film	
Stellar	35mm 2.5 mil base non-perforated
Mapping	5 in., 2.5 mil base non-perforated
Altitude	
Lunar	30 to 80 nautical miles — 50 nm.
Film Capacity	
Stellar	510 feet
Mapping	1500 feet
Cycle Time	
Stellar	8.25 to 33.0 sec/cycle
Mapping	8.25 to 33.0 sec/cycle
Exposure Time	
Stellar	1.5 sec (fixed)
Mapping	1/15 to 1/250 sec
Exposure Control	
(Mapping Only)	Automatic between lens shutter
Forward Motion Compensation	
(Mapping Only)	12.1 to 16.1 milliradian/second (in five discrete steps: 12.1, 13.1, 14.1, 15.1 and 16.1 milliradians/second)

TABLE 2-2 Mapping Camera Characteristics

<b>Resolution</b>	
Stellar	80 lines/mm AWAR, 1000:1 contrast target, EK 3400 film
Mapping	90 lines/mm AWAR, 2:1 contrast target, EK 3404 film
<b>Distortion</b>	
Stellar	±10 microns radial, 5 microns max tangential
Mapping	±50 microns radial, 5 microns max tangential
<b>Overlap</b>	
(Mapping Only)	78 pct (nominal) 67 pct, or 58 pct, adjustable only prior to installation in spacecraft
<b>Camera Generated Recorded Data (see Fig. 2-2)</b>	
<b>Fixed Data</b>	
Stellar	Reseau Fiducial Lens serial number
Mapping	Reseau Fiducial Camera serial number
<b>Auxiliary Data</b>	
Stellar	Index Coded Time Altitude
Mapping	Index Coded Time Altitude Shutter Speed FMC on/off
<b>Power</b>	28 VDC, Operating wattage 100 W min. to 150W max.

TABLE 2-2 CONT'D

- The Aerial Exposure Control (AEC) photocell is pointed in the same direction as the metric lens and "sees" approximately the same surface coverage as the metric camera.
- The Forward Motion Compensation (FMC) drive mechanism provides rectilinear motion of the platen and film in the camera focal plane during exposure to compensate for image motion. The V/h signal can be manually set by the astronaut to five positions by pulsing the Image Motion switch momentarily to the INCR (increase) position (Fig. 1-2). The different V/h rates, accurate to within 3 percent, are: 12.1, 13.1, 14.1, 15.1, and 16.1 milliradians per second. The five step positions and their corresponding altitude ranges are: Step 0, 57-61n.m.; Step 1, 54-57n.m.; Step 2, 54n.m.; Step 3, 65-125 n.m.; Step 4, 61-65 n.m. FMC is disengaged if the switch is in the off position.
- The camera system must be in the STANDBY mode during launch, trans-lunar injection and all other SPSpowered flight phases to prevent slack loops in the film.
- The cameras must be cycled once every 24± 6 hours following film loading to prevent film set. No warmup period is required.
- The system must be turned to the STANDBY mode for warmup 25 hours prior to use for operational photography.
- The nominal lunar altitude is 60 n.m. with an operating range between 40 and 80n.m.
- The cameras are directional and sun sensitive.
- Film flattening in each camera is accomplished by means of a glass focal plane platen and a moveable pressure plate. The emulsion side of the film is in contact with the focal plane platen.

- The laser altimeter is hard mounted to and boresighted with the mapping camera subsystem.
- Both cameras require a supply of gaseous nitrogen (GN2) to provide an inert and pressurized atmosphere to minimize potential static electrical corona discharge which could cause exposed areas on the film. The nitrogen supply is shared with the panoramic camera.
- The exposed film from both cameras is contained in a single recoverable record container that is retrieved via EVA during transearth coast.
- The center of exposure pulse from the mapping camera is used to synchronize the stellar camera (within 20 milliseconds) and laser altimeter (within 10 milliseconds) of the mapping exposure.
- Overlap of the mapping coverage is nominally 78%. This can be adjusted prior to spacecraft installation at discrete steps of 56%, 67%, and 78%.
- The mapping film has fixed data recording including a square array of 121 reseau points, eight fiducial dots and the camera serial number. Auxiliary data recording displays, in binary word form include time, altitude, shutter speed and FMC on/off information (Figs. 2-2 and 2-3). These data are recorded simultaneously with the imagery.
- The stellar film has fixed data recording including a square array of 25 reseau crosses, four edge fiducials and the lens serial number. Auxiliary data, time and altitude are recorded in binary word form at center of exposure time. (Figs. 2-2 and 2-4).

3 February 1971  
 Revised Date:  
 23 July 1971

ROW	17	MINUTE UNITS-BCD	MINUTE TENS-BCD	HOUR UNITS-BCD	HOUR TENS-BCD	UNITS-BCD	DAY TENS BCD	DAY TENS BCD	UNITS-BCD	SECOND UNITS-BCD	SECOND TENS-BCD	UNITS-BCD	SECOND TENS-BCD	CC1_1	CC1_2		
22		X <sub>1</sub>	O <sub>1</sub>	X <sub>2</sub>	X <sub>2</sub>	X <sub>4</sub>	O <sub>1</sub>	O <sub>1</sub>	O <sub>1</sub>	X <sub>2</sub>	X <sub>2</sub>	X <sub>4</sub>	O <sub>1</sub>	◇	×		
23		X <sub>1</sub>	O <sub>1</sub>	X <sub>2</sub>	X <sub>2</sub>	X <sub>4</sub>	O <sub>1</sub>	O <sub>1</sub>	O <sub>1</sub>	X <sub>2</sub>	X <sub>2</sub>	X <sub>4</sub>	O <sub>1</sub>	◇	×		
24		X <sub>1</sub>	O <sub>1</sub>	X <sub>2</sub>	X <sub>2</sub>	X <sub>4</sub>	O <sub>1</sub>	O <sub>1</sub>	O <sub>1</sub>	X <sub>2</sub>	X <sub>2</sub>	X <sub>4</sub>	O <sub>1</sub>	◇	×		
25		X <sub>1</sub>	O <sub>1</sub>	X <sub>2</sub>	X <sub>2</sub>	X <sub>4</sub>	O <sub>1</sub>	O <sub>1</sub>	O <sub>1</sub>	X <sub>2</sub>	X <sub>2</sub>	X <sub>4</sub>	O <sub>1</sub>	◇	×		
26		X <sub>1</sub>	O <sub>1</sub>	X <sub>2</sub>	X <sub>2</sub>	X <sub>4</sub>	O <sub>1</sub>	O <sub>1</sub>	O <sub>1</sub>	X <sub>2</sub>	X <sub>2</sub>	X <sub>4</sub>	O <sub>1</sub>	◇	×		
27		X <sub>1</sub>	O <sub>1</sub>	X <sub>2</sub>	X <sub>2</sub>	X <sub>4</sub>	O <sub>1</sub>	O <sub>1</sub>	O <sub>1</sub>	X <sub>2</sub>	X <sub>2</sub>	X <sub>4</sub>	O <sub>1</sub>	◇	×		
28		X <sub>1</sub>	O <sub>1</sub>	X <sub>2</sub>	X <sub>2</sub>	X <sub>4</sub>	O <sub>1</sub>	O <sub>1</sub>	O <sub>1</sub>	X <sub>2</sub>	X <sub>2</sub>	X <sub>4</sub>	O <sub>1</sub>	◇	×		
29		X <sub>1</sub>	O <sub>1</sub>	X <sub>2</sub>	X <sub>2</sub>	X <sub>4</sub>	O <sub>1</sub>	O <sub>1</sub>	O <sub>1</sub>	X <sub>2</sub>	X <sub>2</sub>	X <sub>4</sub>	O <sub>1</sub>	◇	×		
30		X <sub>1</sub>	O <sub>1</sub>	X <sub>2</sub>	X <sub>2</sub>	X <sub>4</sub>	O <sub>1</sub>	O <sub>1</sub>	O <sub>1</sub>	X <sub>2</sub>	X <sub>2</sub>	X <sub>4</sub>	O <sub>1</sub>	◇	×		
31		X <sub>1</sub>	O <sub>1</sub>	X <sub>2</sub>	X <sub>2</sub>	X <sub>4</sub>	O <sub>1</sub>	O <sub>1</sub>	O <sub>1</sub>	X <sub>2</sub>	X <sub>2</sub>	X <sub>4</sub>	O <sub>1</sub>	◇	×		
32		X <sub>1</sub>	O <sub>1</sub>	X <sub>2</sub>	X <sub>2</sub>	X <sub>4</sub>	O <sub>1</sub>	O <sub>1</sub>	O <sub>1</sub>	X <sub>2</sub>	X <sub>2</sub>	X <sub>4</sub>	O <sub>1</sub>	◇	×		
		TIME WORD								METRIC ONLY							

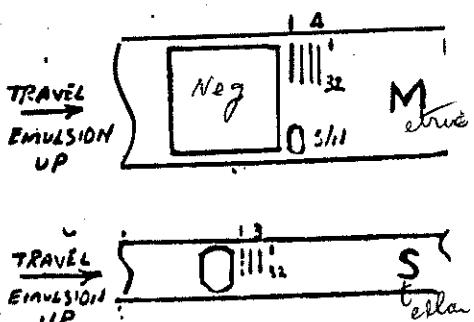
INDEX ROW  
 START OF PASS ROW

### LEGEND

- DOT EVERYTIME  
 BINARY ONE
- BLANK EVERYTIME  
 BINARY ZERO
- ✗ DATA POSITION  
 ONE OR ZERO
- ◇ DON'T CARE POS'N.  
 ONE OR ZERO

INDEX ROW

### MCS ADR



INDEX ROW

Fig. 2-2  
 Auxiliary Data Recording  
 Format

AUX. DATA RECORDING

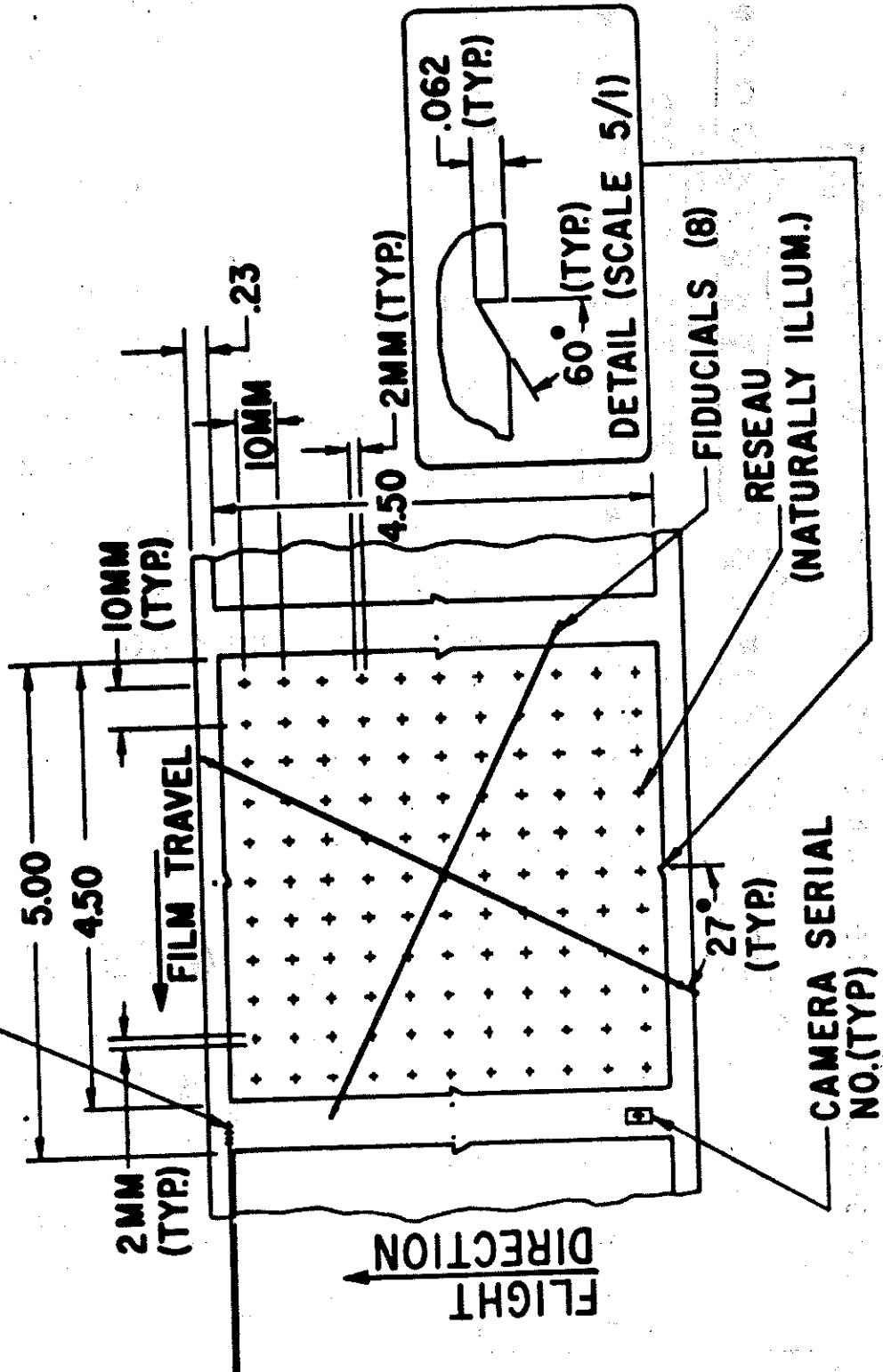


Fig. 2-3 Mapping Camera Film Format

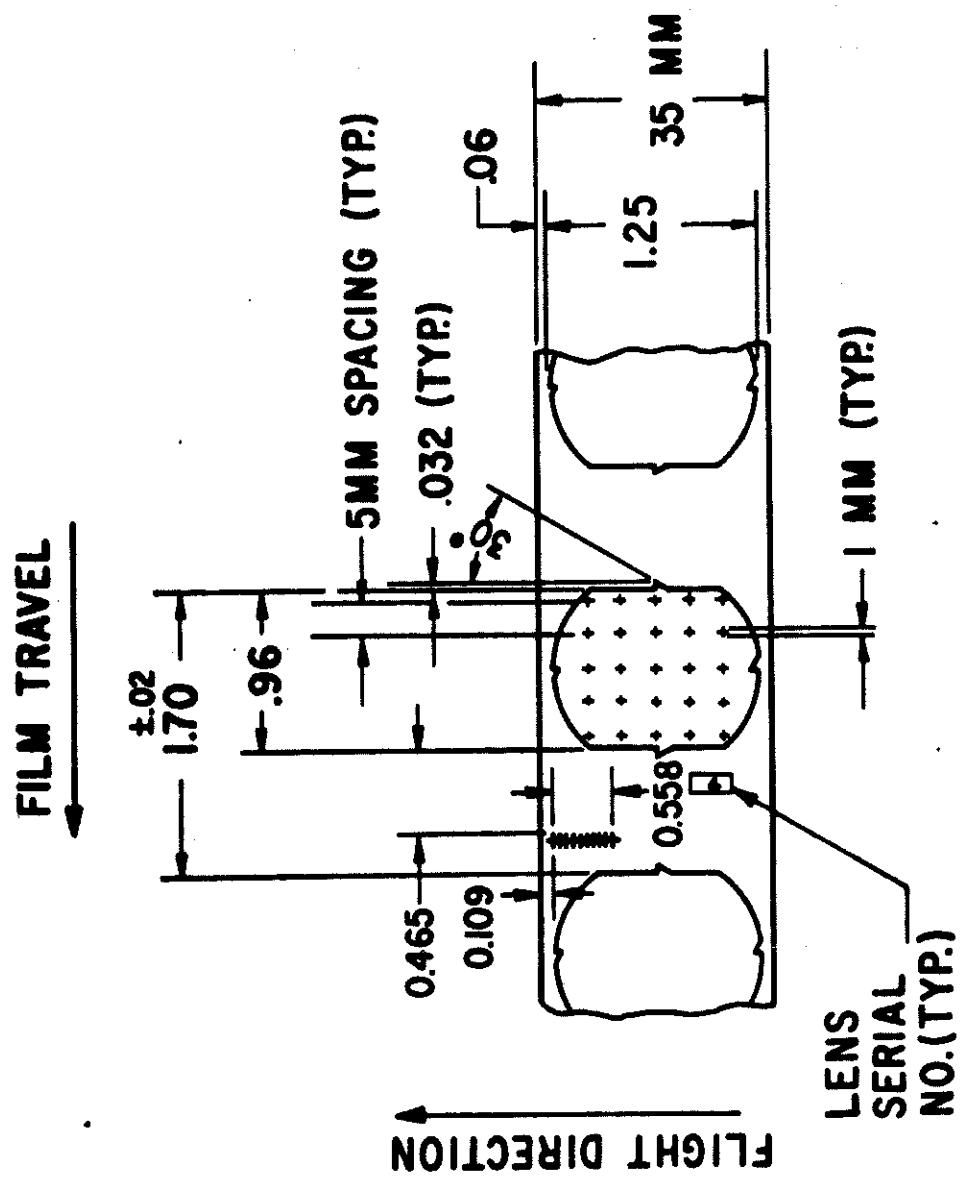


FIG. 2-4 Stellar Camera Film Format

- The mapping camera exposure times are determined by a rotating shutter. Shutter speeds vs. exposure times are as follows:

Shutter Speed (rpm)	Exposure time (sec)
1280	1/250
800	1/156
520	1/101
320	1/62.5
200	1/39
100	1/23.4
80	1/15.6

### 3.0 LASER ALTIMETER

The laser altimeter is an altitude measuring device utilizing a ruby laser and associated optics. The altimeter is hard mounted to the Mapping Camera Subsystem (MCS) and is bore sighted to the mapping camera. It has two modes of operation, Camera and Automatic. In the camera mode, the MCS commands the laser to range with each exposure of the mapping camera. In Automatic the laser automatically ranges once every 20 seconds. For the most part, the altimeter is used in the Camera mode in conjunction with the MCS which provides positional and attitude data. In this mode the altitude is recorded on both the metric and stellar exposures in addition to being telemetered.

During operation the ruby laser and associated optics transmits an intense, short duration, light (laser) pulse to the lunar surface. This pulse is reflected from the surface and is received and detected by the receiving telescope and photomultiplier tube in the altimeter. As the speed of the pulse is known the elapsed time difference between the transmitted and received pulse is easily converted to range by the electronic circuitry.

The output of the ruby laser at the resonant reflector has a diameter of 0.250 inch and a beam width of 4.8 milliradians. A 16-power Galilean telescope expands the transmitted beam to a four-inch diameter with a 0.3 milliradian beamwidth. During the time interval between transmission and reception, the range counter performs the range measurement by counting increments of 6.67 nanoseconds (1 meter) supplied by a master crystal oscillator (149.8962 MHz). The range measurement is then clocked out as a serial word to the spacecraft data system (SDS).

In addition to altitude data, the laser altimeter provides digital and analog equipment status information for telemetering purposes. The digital data consists of 18 bits of range (altitude) data and six bits of equipment status: regulated +5vdc, range overflow, laser output, Q-switch motor status Krytron trigger, and MC data request (Fig. 1-3). The analog data consists of additional equipment status information and includes regulated -5 vdc, photomultiplier voltage, pulse forming network voltage and laser cavity temperature. The operational limits of these are shown in Fig. 1-3

#### 4.0 OPTICAL BAR PANORAMIC CAMERA

The optical bar panoramic camera is configured as a single, self-contained unit with no separately housed components. The camera structure consists of three major assemblies; the roll frame assembly, the gimbal assembly, and the main frame assembly. (Fig. 4-1). The roll frame assembly, which rotates in the scan direction (crosstrack) during camera operation, supports the lens assembly roller cage and variable slit assembly. The roll frame, as well as the light sensor, lens drive, frame roller, and framing roller drive, is supported by the gimbal assembly which tilts fore and aft for stereo coverage and forward motion compensation. The gimbal structure is pivoted on bearings on the main frame along with the gimbal torque motor, V/h sensor, supply and takeup cassettes, film shuttle and the control electronics. The main frame is mounted on rails that are attached to shelves in the SIM with the centerline of the camera displaced 8.25 inches from the centerline of the SIM toward the positive Y coordinate. The optical axis (at center of format) is parallel to the SIM centerline. Basic camera parameters are listed in Table 4-1. Other pertinent facts related to camera installation and operation are listed below.

- The camera assembly envelope is 25 x 29 x 60 inches and weighs 335 pounds including film and cassettes. The film cassette weighs 72 pounds and is 19 inches in diameter and 7 inches wide.
- Since thermal stabilization is variable and dependent upon the SIM thermal environment, a warmup period is required prior to operation for the camera to stabilize at the specified operating temperature ( $88^{\circ} \pm 2^{\circ}\text{F}$ )

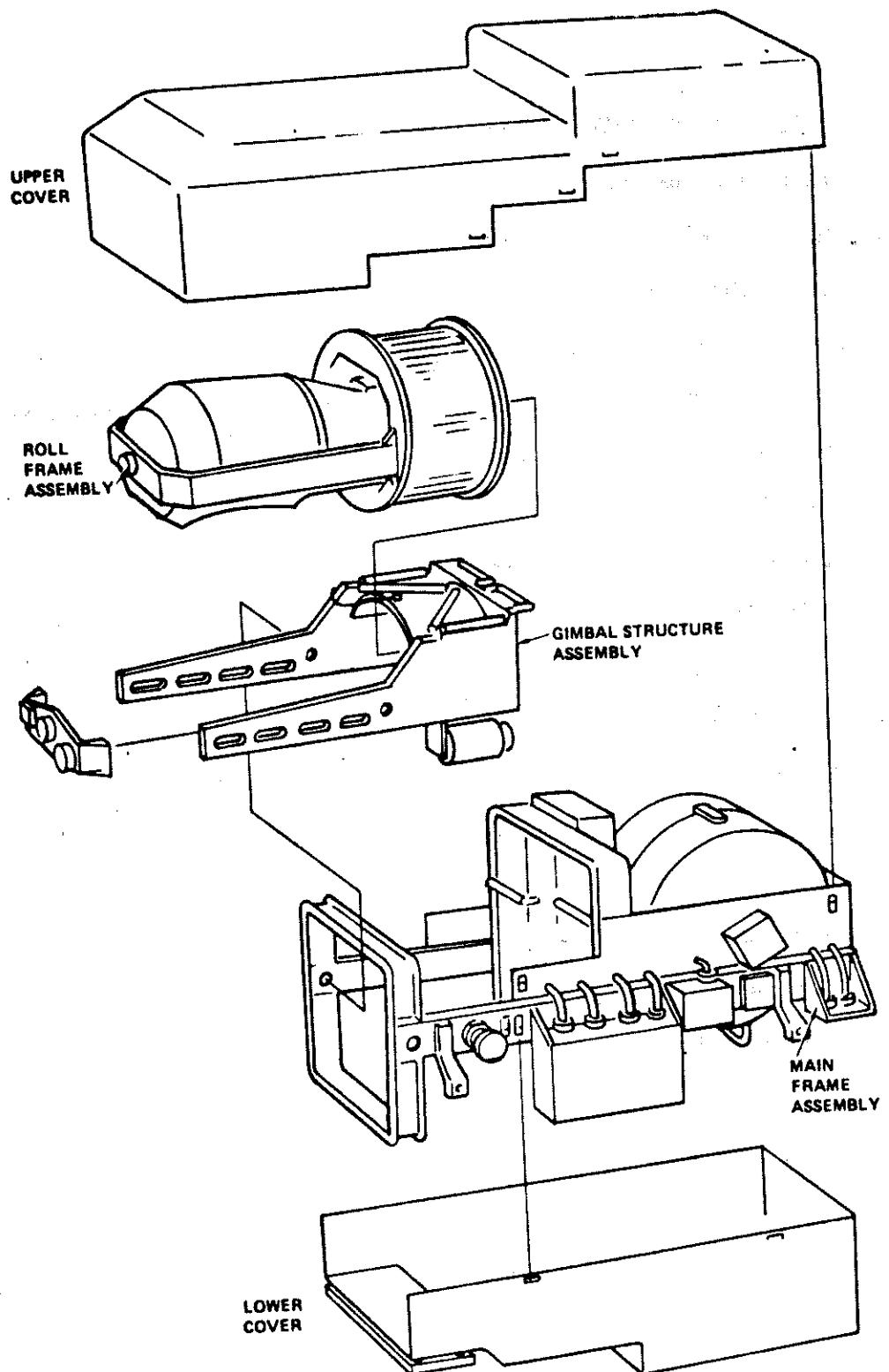


Fig. 4-1 Major Panoramic Camera Assemblies

## CHARACTERISTICS OF THE APOLLO PANORAMIC CAMERA

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Camera	Optical Bar
Lens	
Type	Petzval
Focal Length	24 inches (610 millimeters)
Relative Aperture	f/3.5
Resolution	135 lines per millimeter, 80 percent of targets, no targets less than 108 lines per millimeter, using 2:1 contrast, standard Air Force three- bar targets, 0.002-second exposure with 3404 film, Wratten 23A filter
Field of View	
Along Track	10 degrees, 46 minutes (11.7 nautical miles at 60-nautical-mile altitude) (Fig. 4-5)
Crosstrack scan	108 degrees (183 nautical miles at 60-nautical- mile altitude) (Fig. 4-5)
Overlap	Consecutive forward frames and consecutive aft frames overlap 10 percent; stereo pairs overlap 10 percent.
Shutter	
Type	Scanning slit
Slit Width	0.015 to 0.300 inch
Film	
Type	EK 3414 (6,500 feet; total of 1,650 exposures)
Width	5 inches
Thickness	0.0025 inch
Format	45.24 by 4.5 inches (47.24 inches including data block)
Data (See Fig. 2-5)	Data block for each image frame contains the following information: time (GMT or launch site), data card (shows mission particulars) V/h, frame number, camera pointing attitude (forward, vertical, aft), reference arrows (vertical arrow indicates FMC direction and camera line of flight -- direc- tion of camera line of flight is directly opposite to FMC direction; horizontal arrow points to applicable image frame); fiducial marks are printed on both edges of the film and IRIG B time code is printed on the forward edge of the film.
Exposure Control	Automatic with variable slit
Slit Width Range	0.015 to 0.300 inch
V/h Range	0.010 to 0.019 radian per second
Installation	Parallel to vehicle roll axis

in rad

TABLE 4-1 Characteristics of the Apollo Panoramic Camera

CHARACTERISTICS OF THE APOLLO PANORAMIC CAMERA

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Operating Modes	25-degree convergent stereo at autocycle of 4.7 to 8.9 seconds per cycle; monographic at autocycle of 8.9 to 16.9 seconds per cycle.
Lens Transmission	56%
Telemetry Circuits (See Fig. 1-3)	E (V/h) Exposure Command Slit width Capping Shutter Position Framing roll position Air Solenoid Shuttle position Lens torque current Camera Mechanical Temperature Electronic enclosure temperature Lens (forward) temperature Lens (aft) temperature Lens barrel temperature FMC tachometer

TABLE 4-1 (CONT'D)

- The V/h sensor governs the rotation rate of the roll frame assembly, and FMC rate thus assuring a controlled sequence of imagery related to changes in spacecraft velocity and altitude. Input from the V/h sensor, combined with AEC sensor input, also controls the aperture slit width for correct exposure.

- During camera "operate" or "standby" mode, 14 functional signals can be monitored at ground stations via the vehicle PCM telemetry circuits (Table 4-1).

- The camera must be cycled once every 25 hours  $\pm 6$  hours following flight film loading to prevent film set. This automatically advanced 5 frames, no warmup required.

- The nominal mean lunar altitude of the V/h sensor is set for 60 n.m., thus limiting operation altitude capability between 45 and 80 n.m. to achieve specification quality photography. When operating outside this altitude band, or when the V/h signal is lost, the V/h sensor backup signal commands a nominal V/h value set for a 60n.m. lunar altitude, or 13.6 milliradians per second.

- The pan camera is a directional sensitive instrument (because of FMC) requiring the CSM positive X-axis to be in the direction of the velocity vector during camera operation. The camera is also sun sensitive.

- During pan camera operation the gamma-ray spectrometer boom and mass spectrometer boom must be retracted so as not to be in the optical field-of-view of the camera.

- The exposure can be biased by the astronaut through the Exposure Increase/Decrease switch (Figure 1-2). Placing the switch in Increase,

biases the AEC command to increase exposure by 1/2 f stop. This is accomplished by increasing the size of the slit opening. In the Decrease position the AEC command is biased by 1 f stop by decreasing the slit width.

- Stereographic photography is achieved by alternately tilting the gimbal assembly 12.5° forward and 12.5° aft prior to exposure of each frame. Cycling rate is controlled to insure image overlap between stereo pairs of 10 percent. The stereo-mate to frame number N will be N+5 when frame number N is a forward exposure or N-5 when N is an aft exposure (Fig. 4-2).

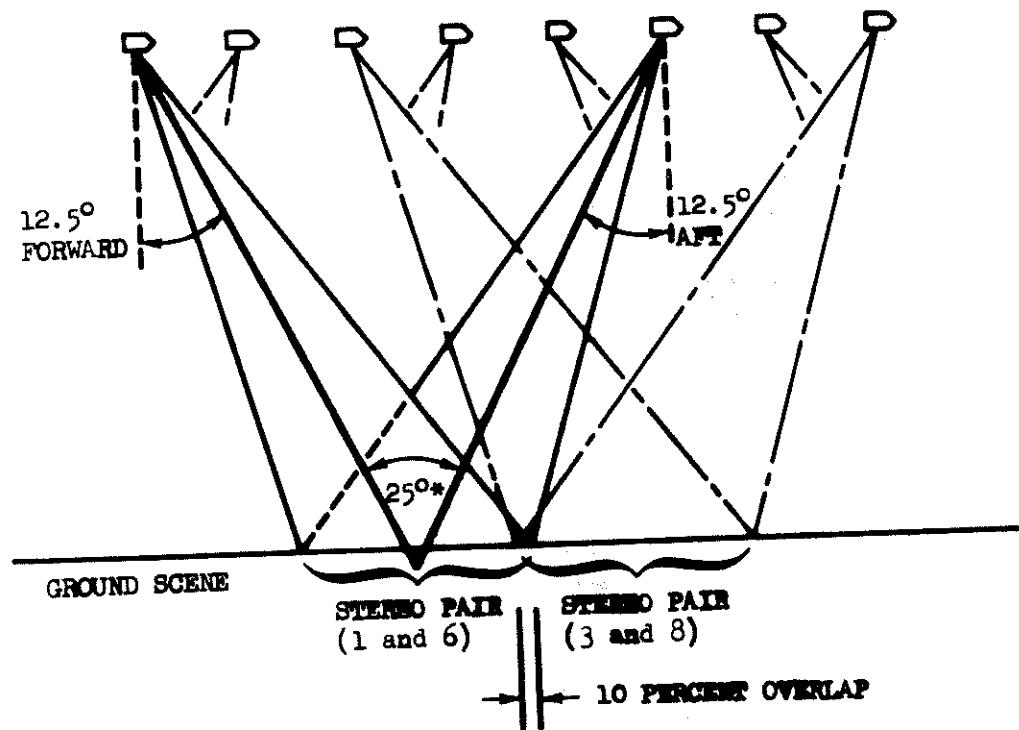
- Scan angle "fiducial" marks are generated along both sides of the photography at 2.5° intervals. The relative angular position of the imaged marks is known to about 5 minutes of arc and the marks are referenced to the center of scan mark at the centerline of the camera (Fig. 4-3).

- IRIG B coded time from the vehicle central timing equipment (CTE) is recorded on the edge of the film. The complete time word is recorded every second in seconds, minutes, hours and days for the instant of exposure of the reference mark. Timing marks are provided continuously in 10 millisecond increments between successive time words (Figs. 4-3 and 4-4). Spacing between timing marks as recorded on the film will vary directly with the cycle rate (V/h rate) of the camera.

- Operating time is limited to 30 minutes per pass to prevent overheating of the camera.

- The post-flight pointing knowledge requirement of the pan camera is  $\pm 2^\circ$  (three axes) thus requiring a spacecraft control deadband of  $\pm 0.5^\circ$  during operation. The root sum square (RSS) of the altitude uncertainty is  $\pm 1.44^\circ$ .

b2  
Careful  
115  
Sweeping  
Constant  
Scan rate



**\*CONVERGENCE ANGLE**

Fig. 4-2 - Stereo coverage and overlap (25-degree convergence angle)

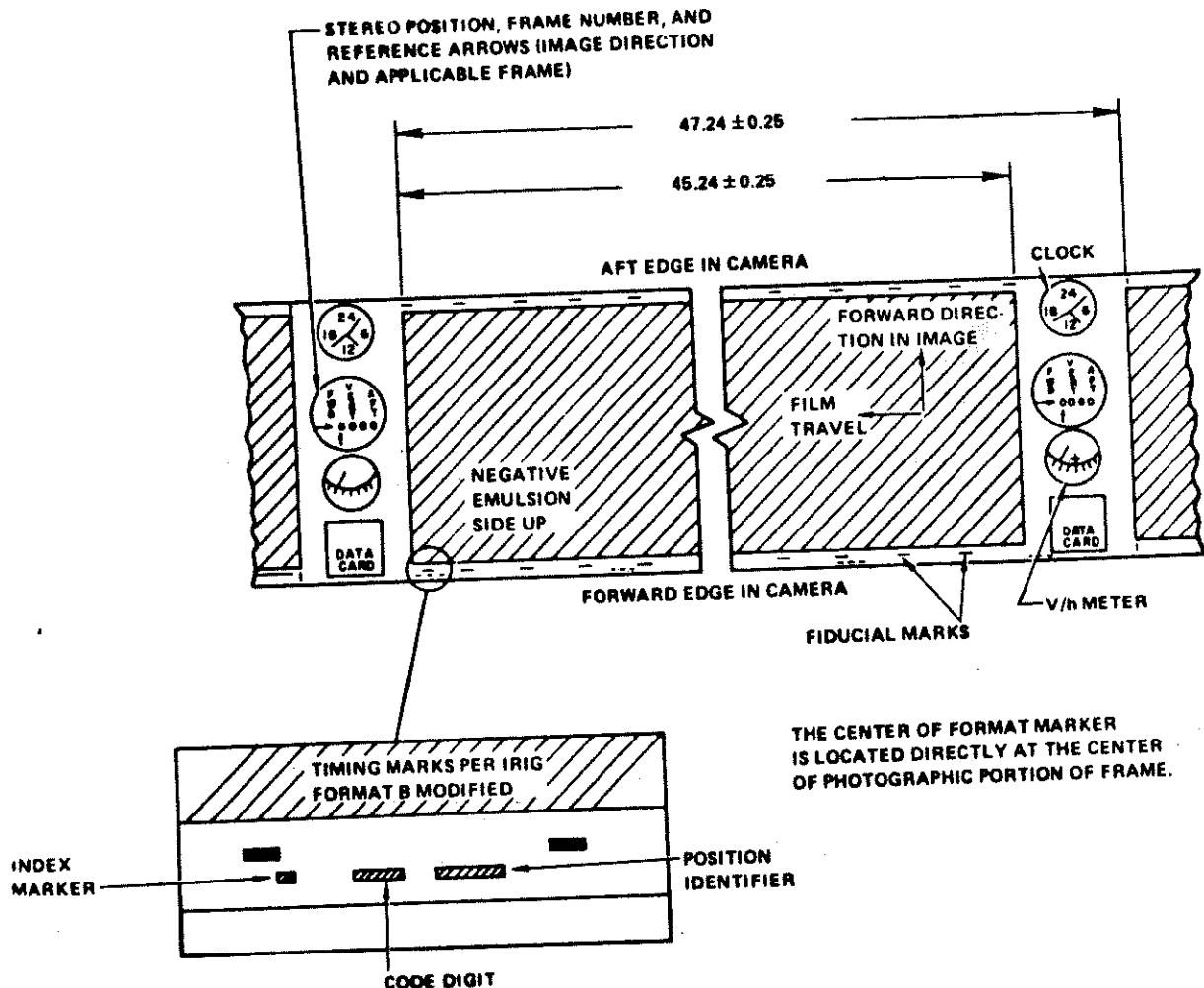


Fig. 4-3 Panoramic Camera Film Format

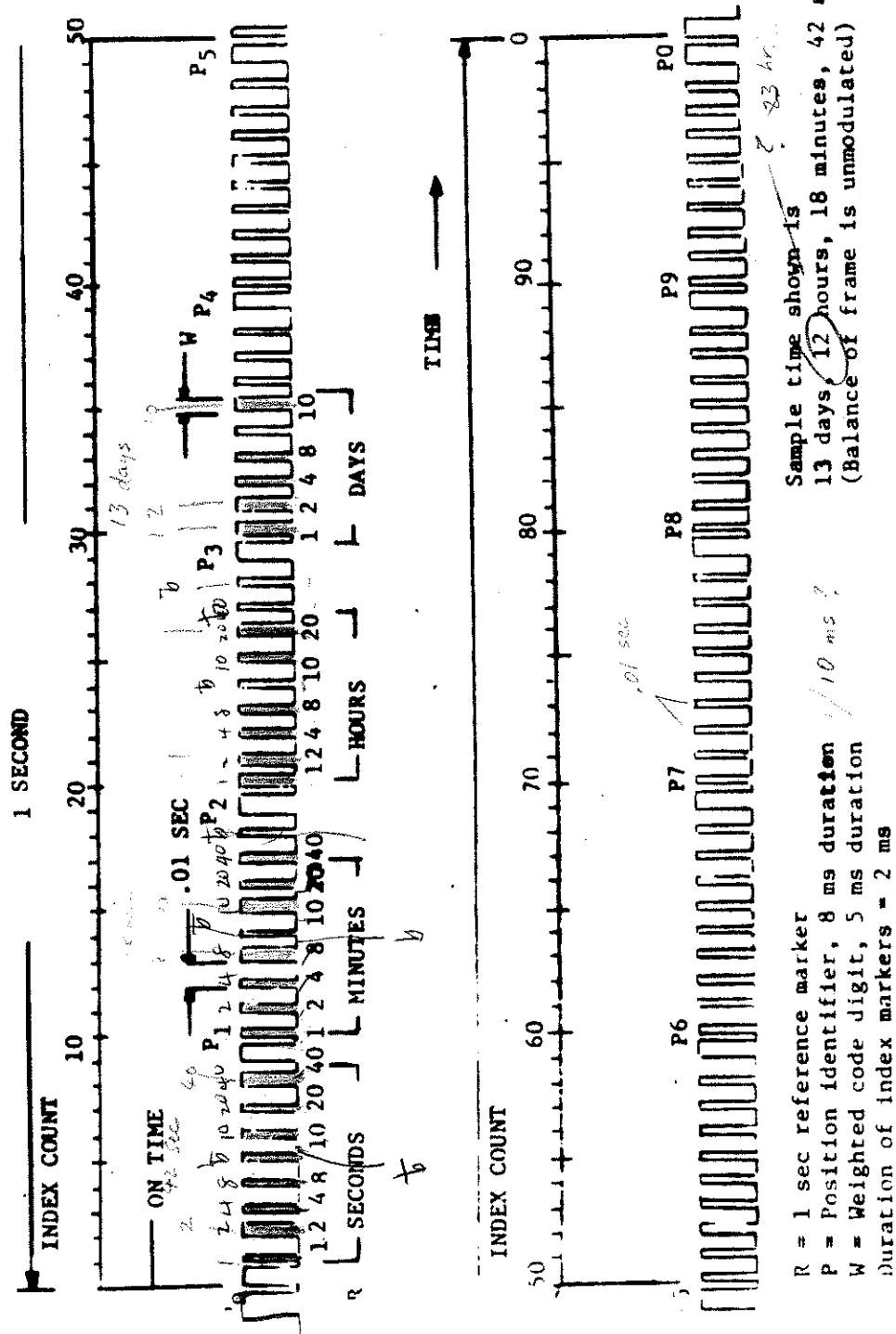


Fig 4-4 100 PPS Code - Modified IRIG Standard Format B

- An astronaut will retrieve the film cassette via EVA during transearth coast.
- This experiment does not require concurrent performance of any other experiment; however, it is desirable the mapping camera be operated during panoramic camera operation.

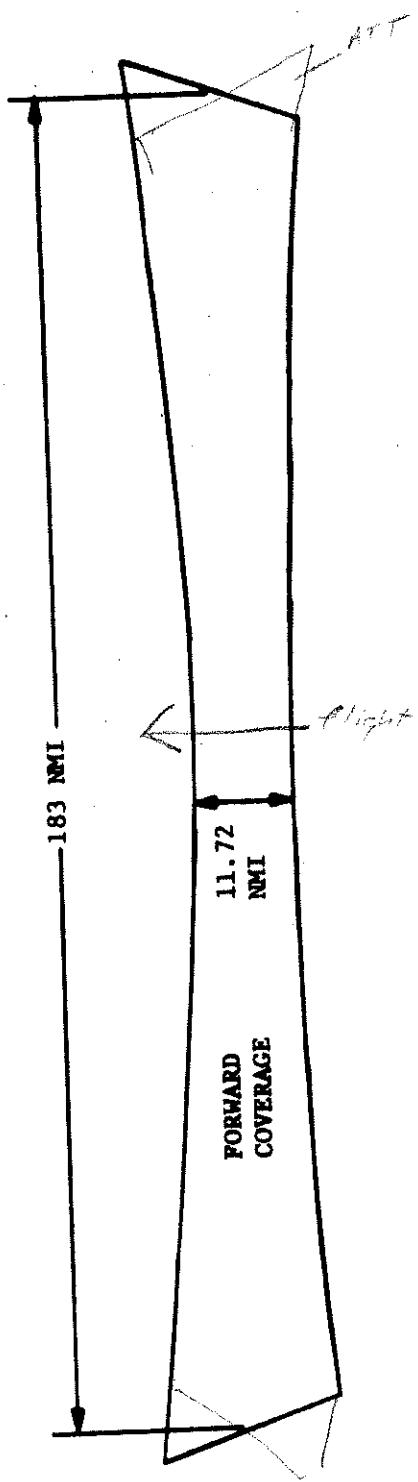


Fig. 4-5 LUNAR SURFACE FOOTPRINT FOR 60 NMi ORBIT

## 5.0 APOLLO ORBITAL SUPPORT DATA (PHOTOGRAPHIC EPHEMERIS)

The Apollo Photograph Evaluation (APE) program is a UNIVAC 1108 computer program developed with the primary function to generate camera position, orientation and evaluation data for use in the photographic phases of Apollo missions. The APE program links together a spacecraft attitude history and camera attitude history to form a data base for its computations. APE operates on these data by way of a generalized camera model to:

1. Define the camera orientation
2. Define the aiming direction of the camera optical axis.
3. Locate on the surface of the central body
  - a. the spacecraft nadir
  - b. the principal intersection point (intersection of the camera optical axis direction with the central body surface)
  - c. points along the boundary of the photograph footprint on the central body surface.
4. Define geometrical relationships between the local vertical to the central body surface and the camera axes coordinate system.
5. Define lighting parameters at the principal point of intersection.
6. Define first-order uncertainties in selected location and camera orientation parameters.

The APE program is configured to operate on the following input data combination.

1. A Houston Operations Predictor/Estimator (HOPE) generated moon-centered, CSM trajectory with state vectors and a covariance matrix defined

at times coinciding with actual camera shutter times (Downlinked camera shutter times are input to HOPE through APE). (Tape Input)

2. Actual telemetered histories of the spacecraft gimbal angles.

(Tape Input).

3. Histories of the camera positioning angles, input via cards, except for those for the sextant camera (shaft and trunnion angles) which are input along with the vehicle gimbal angles on the input telemetered data tape.

4. The JPL DE19 planetary ephemeris for stellar camera only. (Tape input).

5. The SAO star catalog for stellar camera only (Tape Input).

Operating on the above combination of inputs, APE computes for all except the stellar camera:

1. The rotation angles ( $\phi, \omega, \kappa$ ) which reference the camera axis coordinate system to the local horizontal system.

2. The camera optical axis tilt ( $t$ ). The angle between the local vertical and the camera optical axis.

3. Tilt azimuth ( $\alpha$ ). The angle between north and the projection of the camera optical axis on the local horizontal plane.

4. X-tilt ( $t_x$ ) and y-tilt ( $t_y$ ). The angles between the local horizontal plane and the camera x and y axes respectively.

5. Heading ( $H$ ). The angle between north and the projection of the camera x-axis on the local horizontal plane.

6. Swing ( $S$ ). The angle at the principal point of intersection between the positive camera y-axis and the nadir point.

7. The latitude ( $\Delta$ ) and longitude ( $\lambda$ ) of nadir taken directly from the input trajectory.

8. The latitude ( $\delta$ ) and longitude ( $\eta$ ) of the principal point of intersection.

9. At the principal point of intersection, the sun elevation ( $\mu$ ) azimuth ( $\sigma$ ) and phase ( $\rho$ ) angles.

10. The latitude and longitude of points along the boundaries of the photo footprint on the lunar surface.

Employment of APE in the star pattern prediction mode will allow for the following operations:

1. Computation of the direction cosine matrix relating to the stellar camera diapositive coordinate system to the geocentric mean of 1950 coordinate system.

2. Definition of the stellar camera optical axis aiming direction in terms of mean of 1950 right ascension ( $\alpha$ ) and declination ( $\delta$ ).

Specifically, the basis for all APE camera orientation and related definitions is the APE-constructed generalized direction cosine matrix [M]. This is the product of transformation matrices which defines the relationship between the camera axes coordinate system and the local horizontal coordinate system. This matrix derives its generality with regards to various camera systems through its structure of the matrix component which expresses the transformation from the spacecraft body axes coordinate system to the camera axes coordinate system. It derives its generality with regards to various central bodies through its matrix component which expresses the transformation from the local horizontal system to the mean of nearest Besselian year coordinate system.

The direction cosine matrix employed for processing CSM lunar orbital photography is:

$$[M] = [A] [B] [C] [D] \quad (1)$$

where

$$[A] = [\rho]_z [3\pi/2 + \gamma]_y [\delta]_z; (\rho, \gamma, \delta = \text{camera mounting angles}) \quad (2)$$

= Transformation matrix from the spacecraft body axes coordinate system to the camera axes coordinate system.

$$[B] = [\psi]_x [\theta]_z [\beta]_y \quad (\psi, \theta, \beta = \text{spacecraft gimbal angles}) \quad (3)$$

= Transformation matrix from the spacecraft platform (IMU) coordinate system to the spacecraft body axes coordinate system

$$[C] = [\text{REFSMMAT}]$$

= Transformation from the mean of nearest Besselian year (NBY) coordinate system to the spacecraft platform coordinate system

= [E] [F]; in which,

[E] = Transformation matrix from the selenographic coordinate system to the mean of nearest Besselian year coordinate system.

[F] = Transformation matrix from the local horizontal coordinate system to the selenographic coordinate system.

The characteristics of these matrix elements of [M] are as follows:

- [A]

As a triple rotation, [A] is adaptable to any spacecraft-mounted experiment. Its utility is realized by assigning proper values to  $\rho$ ,  $\gamma$  and  $\delta$  as determined from the decomposition of the various spacecraft body axes to the camera axes transformations into  $\rho$ ,  $\gamma$  and  $\delta$  equivalents. The assigned mounting angles are preflight determined and are input to APE via cards.

- [B]

The spacecraft gimbal angle values necessary for the construction of this matrix are provided to APE by the input telemetered data tape. The values employed by APE are interpolated values determined by a two parabola averaging technique.

- [C]

This matrix [REFSMMAT] is a card input to APE, whose values are determined in the RTCC.

- [E]

This matrix is determined by way of the standard TRW rotation routine ROTAT which is employed in all Houston Operations trajectory programs.

- [F]

This transformation matrix is developed in APE via a set of vector equations. The values of its elements are determined from the input (trajectory tape) spacecraft coordinates expressed in the body-centered inertial coordinates of the central body (selenographic for the moon) and the card input central body mean radius values.

When APE is operating under the star pattern prediction for the stellar camera it is used to:

1. Define the orientation of the camera axes coordinate system relative to the mean of 1950 coordinate system.
2. Define the camera optical axis aiming direction in terms of the star catalog star location parameters (mean of 1950 right ascension and declination).

3. Define the stellar camera field of view bounds in terms of these same parameters

4. Select from the input star catalog all stars within a specified brightness range which lie within these bounds.

5. Select all planets from the input planetary ephemeris that lie within the field of view bounds.

6. Correct the star catalog positions of the selected stars for proper motion and aberration.

7. Correct the positions of the selected planets for aberrations.

8. Compute and plot the rectangular coordinates of the projections of selected star and planet positions onto the stellar camera diapositive.

When serving this purpose APE defines camera axes orientations via an [M] relating the stellar camera diapositive coordinate system to the mean of nearest Besselian year coordinate system. Computations for definition of stellar camera field of view boundaries and of stellar camera diapositive star coordinates are substituted for the normal surface location computations and star catalog search and select logic is added.

Sample printouts of the ephemeris for the mapping and panoramic cameras are shown in Figures 5-1,5-2. The laser altimeter altitude for each mapping camera exposure is included in the mapping camera ephemeris. Those familiar with the Lunar Orbiter experiment will note that the format is similar in composition and content. As stated earlier the 16mm microfilm of the orbital support data will not be available for several months after the mission.

APOLLO 12 840,41 SS2 METRIC/70 FIN FRAME - 1

	YEAR	MONTH	DAY	HOUR	MINUTE	SECOND	
GMT	69	11	21	9	56	30.000	
GET			6	17	34	30.000	
STATE VECTOR	X			Z	X	Y DOT	Z DOT
1950.0 U	-1468.967697	-902.6637192	-647.3101807	-0.9821931	-1.1099241	-0.6822385	-0.2115187
SELDINGRAPHIS	1771.41119413	423.5657168	-272.7088356	-0.4111940	-1.5670300	.0000	.0000
SIGMA AISELNS	.00						
LONGITUDE OF NADIR POINT	13.4476075	DEG	LATITUDE OF NADIR POINT				-8.5155866 DEG
SIGMA NADIR LONGITUDE	.0000000	DEG	SIGMA NADIR LATITUDE				.0000000 DEG
LON OF CAMERA AXIS INTERSECT	14.1854523	DEG	SIGMA NADIR AXIS INTERSECT				-8.5280169 DEG
SPACECRAFT ALTITUDE	1841.6509552	KM	LATI OF CAMERA AXIS				103.5609283 KM
SIGMA SPACECRAFT VELOCITY	.0000000	KM	SPACECRAFT ALTITUDE				277.5374146 DEG
MEAN ALTITUDE RATE	.0003600	KM/SEC	AZIMUTH OF VELOCITY				1.63345675 KM/SEC
TILT AZIMUTH	91.0305700	DEG	SIGMA TILT ANGLE				12.0504202 DEG
SIGMA TILT AZIMUTH	5.4383475	DEG	SIGMA TILT ANGLE				4.3532313 DEG
SUN ELEVATION AT PRIN GND PNT	67.4644011	DEG	SUN AZIMUTH AT PRINCIPAL GRND PNT				72.9538734 DEG
LONGITUDE OF SUBSOLAR POINT	35.6841515	DEG	LATITUDE OF SUBSOLAR POINT				-1.48246430 DEG
ALPHA	-12.5134517	DEG	SWING ANGLE				82.6532001 DEG
EMISSION ANGLE	12.7804220	DEG	SIGMA SWING ANGLE				2.1758550 DEG
PHASE ANGLE	34.9021535	DEG	NORTH DEVIATION ANGLE				171.7547817 DEG
PHI	-11.9545481	DEG	A-TILT				1.5298157 DEG
SIGMA PHI	.7333002	DEG	SIGMA X-TILT				.5311160 DEG
KAPPA	-171.7829197	DEG	Y-TILT				1.9502243 DEG
SIGMA KAPPA	.2274052	DEG	SIGMA Y-TILT				.7145555 DEG
OMEGA	1.5298157	DEG	HEADING				-0.4589548 DEG
SIGMA OMEGA	.5311160	DEG	SIGMA HEADING				.7807548 DEG
SCALE FACTOR	.00047544	M/KM	LASER SLANT RANGE				109.0000000 KM
SPACECRAFT ALTITUDE(LASEN)	103.0000000	KM	ALTITUDE DIFFERENCE				.5609283 KM

FIG. 5-1 Metric Orbital Support Data

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MAGNITUDE (KM) 106.038651  
 Selenographic Direction Cosines X -0.98974986 Y -0.02203455 Z 14110170  
 of Camera Axis

## FRAME CORNER POSITIONS

LONG 120.8303 DEG LONG 130.1993 DEG  
 LATI +90.4662 DEG LATI -70.2419 DEG

\* 68.14852 KM (1)  
 (1)

75.0423405 KM

76.088946 KM

TRANSFORMATION MATRIX FROM  
SELENOCENTRIC TO CAMERA

-0.96747792+00	+14529921+00	-0.20706184+00
-0.14287485+00	-0.98938063+00	-0.26697149-01
-0.20874205+00	+0.37550271-02	+0.97796351+00

(3)  
 \* 79.070872 KM  
 (4)  
 \*  
 LONG 150.3378 DEG LONG 150.7499 DEG  
 LATI -10.0003 DEG LATI -70.4015 DEG

TRANSFORMATION MATRIX FROM  
LOCAL HORIZONTAL TO CAMERA

-0.42445775+00	+75740624+00	+0.49615661+00
+0.40231946-01	+0.56610373+00	-0.82292177+00
-0.90416220+00	-0.32536492+00	-0.27681629+00

FIG. 5-1 (CONT'D)

APOLLO 12 R40 PAN 9/70 FIN FRAME - 201, RELATED FRAME 1

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	YEAR	MONTH	DAY	HOUR	MINUTE	SECOND	
GMT GET	69 6	11 6	21 17	9 34	56 30	000 000	
STATE VECTOR X	-902.6637192	-647.3101807	-987.1831	A DUT	Y DOT	Z DOT	
1950.0	-1468.9674971	-272.7088356	-1111.940		1.099243	0.6822385	
SELENOGRAPHIC	423.5657768	989.91	6.509		-1.5676300	0.2115167	
SIGMA(SELENO)	7337.44				1.557	1.002	
LONGITUDE OF NAIR POINT	13.04476075	DEG	LATITUDE OF NADIR POINT	+6.5155366	DEG		
SIGMA NAIR LONGITUDE	4.1636403	DEG	SIGMA NADIR LATITUDE	.5445728	DEG		
LONG OF CAMERA AXIS INTERSECT	12.7193147	DEG	LATT OF CAMERA AXIS INTERSECT	.3492838	DEG		
SPACECRAFT RADIUS	1841.6509552	KM	SPACECRAFT ALTITUDE	103.5609283	KM		
SIGMA SPACECRAFT RADIUS	0026422	KM	AZIMUTH OF VELOCITY VECTOR	277.5379146	DEG		
MEAN ALTITUDE RATE	0036009	KM/SEC	HORIZONTAL VELOCITY	1.6345675	KM/S		
TILT AZIMUTH	282.9449196	DEG	TILT ANGLE	12.2033944	DEG		
SIGMA TILT AZIMUTH	0000000	DEG	SIGMA TILT ANGLE	0.0000000	DEG		
SUN ELEVATION AT PRIN GRND PNT	66.1241274	DEG	SUN AZIMUTH AT PRINCIPAL GRND PNT	74.5166362	DEG		
LONGITUDE OF SUBSOLAR POINT	35.6861515	DEG	LATITUDE OF SUBSOLAR POINT	-1.4626630	DEG		
ALPHA	7.7377850	DEG	SWING ANGLE	275.1296065	DEG		
EMISSION ANGLE	12.9427638	DEG	SIGMA SWING ANGLE	*0000000	DEG		
PHASE ANGLE	13.08712469	DEG	NORTH DEVIATION ANGLE	172.3967255	DEG		
PHI	12.1559746	DEG	X-TILT	1.0629695	DEG		
SIGMA PHI	0000000	DEG	SIGMA X-TILT	*0000000	DEG		
KAPPA	-172.0695686	DEG	Y-TILT	-12.1537701	DEG		
SIGMA KAPPA	0000000	DEG	SIGMA Y-TILT	*0000000	DEG		
OMEGA	1.0829695	DEG	HEADING	-82.3023259	DEG		
SIGMA OMEGA	0000000	DEG	SIGMA HEADING	*0000000	DEG		
SCALE FACTOR	00007540	M/KM					

FIG. 5-2 Panoramic Orbital Support Data

S E L E N O C R A P H I C D I R E C T I O N C O S I N E S      X    - . 9 8 9 7 4 4 9 8 6  
O F C A M E R A A X I S      Y    - . 0 2 2 0 3 4 5 5  
Z    + 1 4 1 1 0 1 7 0

F R A M E C O R N E R P O S I T I O N S  
L O N G    1 2 ° 0 . 6 3 0 3 D E G    L O N G    1 3 ° 1 9 9 3 D E G  
L A T I    + 9 ° 4 4 5 6 2 D E G    L A T I    - 7 ° 2 4 1 9 D E G

( 1 )      6 8 ° 1 4 8 5 2 K M      ( 2 )

1 5 ° 4 3 4 6 3 K M

7 6 ° 8 8 7 4 6 K M

M A G N I T U D E ( K M )  
1 0 6 . 0 3 8 6 5 1

T R A N S F O R M A T I O N M A T R I X F R O M  
L O C A L H O R I Z O N T A L T O C A M E R A

- . 9 6 7 4 7 7 9 2 + 0 0      + 1 4 5 2 9 9 2 1 + 0 0      - . 2 0 7 0 6 1 8 4 + 0 0  
- . 1 4 2 8 7 4 8 5 + 0 0      - . 9 8 9 3 8 0 6 3 + 0 0      - . 2 6 6 9 7 1 4 9 + 0 1  
- . 2 0 8 7 4 2 0 5 + 0 0      + 3 7 5 5 0 2 7 1 - 0 2      + 9 7 7 9 6 3 5 1 + 0 0

T R A N S F O R M A T I O N M A T R I X F R O M

S E L E N O C E N T R I C T O C A M E R A

+ 4 2 4 4 5 7 7 5 + 0 0      + 7 5 7 4 0 6 2 4 + 0 0      + 4 9 6 1 5 6 6 1 + 0 0  
( 3 )      + 1 8 2 3 1 9 4 6 - 0 1      + 5 6 6 1 3 3 7 3 + 0 0      + 8 2 2 9 2 1 7 7 + 0 0  
( 4 )      - . 9 0 4 1 6 2 2 0 + 0 0      - . 3 2 5 3 6 4 9 2 + 0 0      - . 2 7 6 8 1 8 2 9 + 0 0

L O N G    1 2 ° 3 3 7 8 D E G    L O N G    1 5 ° 7 4 9 9 D E G  
L A T I    - 1 0 ° 0 0 0 3 D E G    L A T I    - 7 ° 4 0 1 5 D E G

FIG. 5-2 (Con'td)

## 6.0 CAMERA CALIBRATION DATA

Calibration of the Apollo 15 SIM bay cameras was provided by the equipment manufacturers. Copies of the original calibration reports are presented in this section. Although portions of the data may be difficult to read, it seemed more appropriate to include the original reports than to have them retyped. No additional calibration was accomplished by NASA/MSC.

### 6.1 CALIBRATION DATA FOR THE MAPPING CAMERA SUBSYSTEM (MCS)

The stellar calibration report was prepared by Raytheon/Autometric under contract to Fairchild Space and Defense System. In addition to the camera calibrations the report includes calibrated values of the reseaus and the relative orientation angle between the connected lens cones. Fairchild data concerning relative illumination and resolving power of the cameras are included.

TR-71-3404-2R

FINAL

CAMERA CALIBRATION REPORT

CAMERA UNIT SN-003

JULY 1971

Prepared by

Autometric Operation  
Equipment Division  
Raytheon Company

4217 Wheeler Avenue  
Alexandria, Virginia 22304

Prepared for

Fairchild Space and Defense Systems  
300 Robbins Lane  
Syosset, L.I., New York 11791

FOREWARD

This final stellar calibration report is submitted to NASA Manned Spacecraft Center, Houston, Texas, by Fairchild Space and Defense Systems (FSDS), A Division of Fairchild Camera and Instrument Corporation under NASA Contract No. NAS 9-10457. This report was prepared by Raytheon/Autometric, Alexandria, Virginia, under separate contract to FSDS. It supersedes the preliminary version which was previously submitted in May 1971.

This report summarizes the results of the stellar calibration of Mapping Camera Subsystem (MCS) Serial Number 003. The actual stellar field calibration was performed at the NASA White Sands Test Facility (WSTF), Las Cruces, New Mexico, during the period 25-26 March 1971. The resulting stellar imagery was subsequently reduced by Raytheon/Autometric and computer processed and outputted into the form presented herein. It should be emphasized that the calibration values shown for the individual terrain lens/stellar lens internal geometry parameters and relative orientation between connected lens cones are indicative of the capabilities of MCS 003, as derived by stellar methods. Also included in this report are the calibrated values of the terrain and stellar reseau coordinate intersections.

All calibration values presented herein should be utilized in the subsequent post-flight data reductions of MCS 003 terrain and stellar imagery.

A. D. Beccasio  
A. D. Beccasio  
Senior Staff Engineer

Approved by: A. G. Hutchins  
A. G. Hutchins  
MCS Program Manager

12 July 1971  
Fairchild Space & Defense Systems  
300 Robbins Lane  
Syosset, New York 11791

CAMERA CALIBRATION RESULTS

Lunar Mapping Camera SN-003 Stellar Calibration

Terrain Lens (202) Constants of Internal Geometry

$$\begin{array}{ll} EFL = 76.054 \text{ mm.} & S.D. = 0.002 \text{ mm.} \\ CFL = 76.080 \text{ mm.} & S.D. = 0.002 \text{ mm.} \end{array}$$

Principal Point With Respect to Indicated Principal Point  
 (Indicated principal point  $x_{ipp} = 0.0 \text{ mm.}, y_{ipp} = 0.0 \text{ mm.}$ ),

$$\begin{array}{ll} x_p = -0.006 \text{ mm.} & S.D. = 0.001 \text{ mm.} \\ y_p = -0.002 \text{ mm.} & S.D. = 0.001 \text{ mm.} \end{array}$$

## Radial Distortion Parameters

$$\begin{array}{ll} K_1 = -0.13361854 \times 10^{-5} & S.D. = 0.39272852 \times 10^{-7} \\ K_2 = 0.52261757 \times 10^{-9} & S.D. = 0.15933971 \times 10^{-10} \\ K_3 = -0.50728336 \times 10^{-13} & S.D. = 0.19335059 \times 10^{-14} \end{array}$$

## Lens Decentration Distortion Parameters

$$\begin{array}{ll} J_1 = -0.54958195 \times 10^{-6} & S.D. = 0.48922291 \times 10^{-6} \\ J_2 = -0.46089420 \times 10^{-10} & S.D. = 0.11215963 \times 10^{-9} \\ \phi_0 = 2.9659070 \text{ radians} & S.D. = 0.23946572 \end{array}$$

Stellar Lens (104) Constants of Internal Geometry

EFL = 75.713 mm. S.D. = 0.011 mm.  
 CFL = 75.735 mm. S.D. = 0.011 mm.

Principal Point With Respect to Indicated Principal Point  
 (Indicated principal point  $x_{ipp} = 0.0$  mm.,  $y_{ipp} = 0.0$  mm.),

$x_p = 0.042$  mm. S.D. = 0.019 mm.  
 $y_p = -0.003$  mm. S.D. = 0.018 mm.

Radial Distortion Parameters

$K_1 = -0.92376425 \times 10^{-5}$	S.D. = $0.40956407 \times 10^{-5}$
$K_2 = 0.53810549 \times 10^{-7}$	S.D. = $0.33179901 \times 10^{-7}$
$K_3 = -0.70032005 \times 10^{-10}$	S.D. = $0.80049661 \times 10^{-10}$

Lens Decentration Distortion Parameters

$J_1 = 0.81424290 \times 10^{-6}$	S.D. = $0.10033483 \times 10^{-4}$
$J_2 = -0.20686782 \times 10^{-7}$	S.D. = $0.42974779 \times 10^{-7}$
$\phi_0 = 1.2355515$ radians	S.D. = 0.75665362

Results of Lock-Angle Calibration

Relative Orientation Matrix Defining a Transformation from the Terrain Camera to the Stellar Camera

0.999999850	-0.000517921	0.000180754
0.000125645	-0.104494099	-0.994525499
0.000533973	0.994525372	-0.104494018

Relative Orientation Angles (Degrees, Minutes, Seconds)

OMEGA = -95 59 52.859
PHI = 0 1 50.140
KAPPA = -0 0 25.916

Covariance Matrix

$0.98562 \times 10^{-11}$	$-0.33758 \times 10^{-12}$	$0.88273 \times 10^{-12}$
$-0.33758 \times 10^{-12}$	$0.22323 \times 10^{-10}$	$0.34486 \times 10^{-12}$
$0.88273 \times 10^{-12}$	$0.34486 \times 10^{-12}$	$0.35148 \times 10^{-9}$

Standard Deviation of Orientation Angles (Arc-seconds)

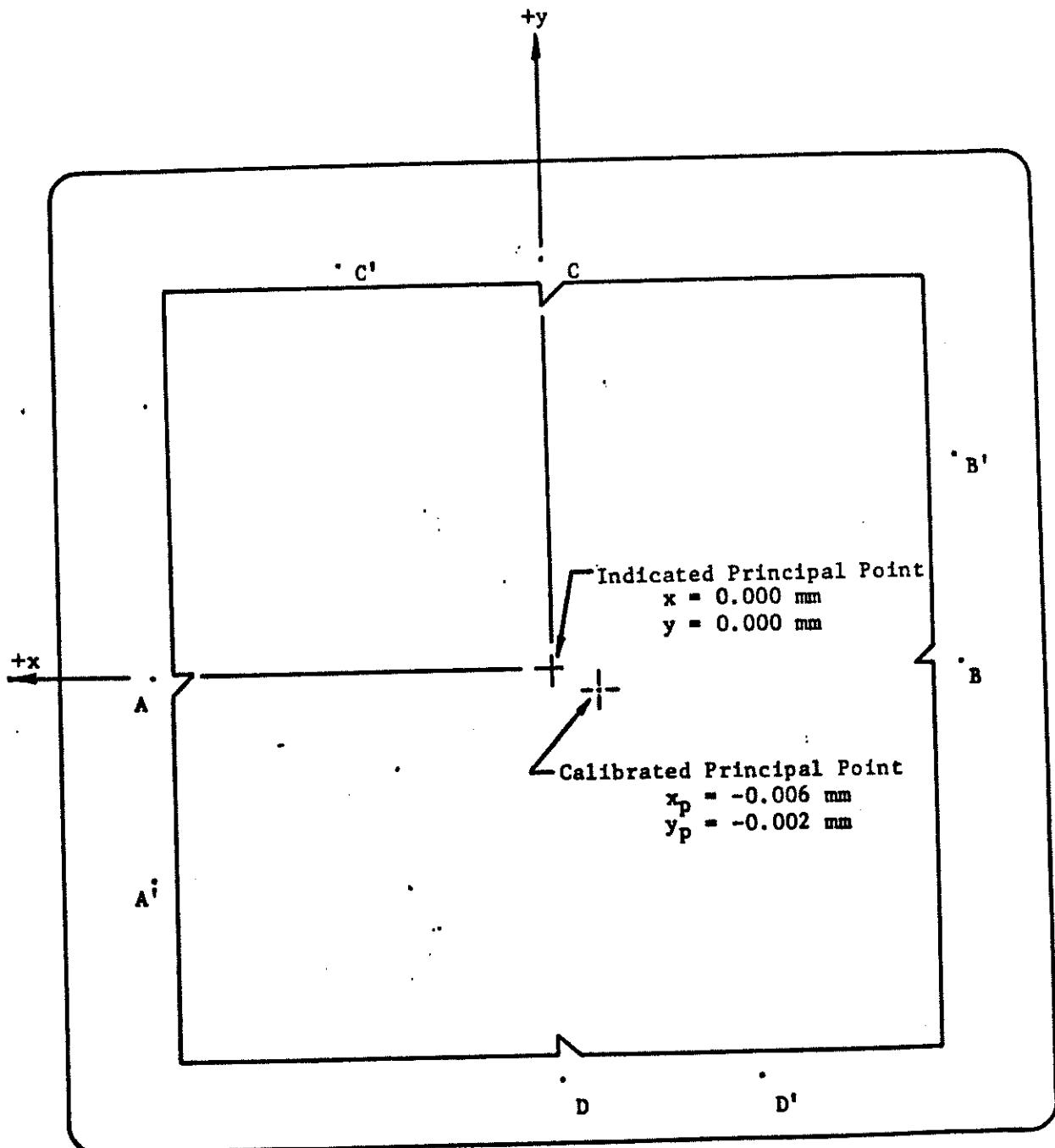
S.D. OMEGA = 0.65
S.D. PHI = 0.98
S.D. KAPPA = 3.87

Statistical Data From Simultaneous Solution

Weighted Sum of Squares = 0.031920

Degrees of Freedom = 1996

Standard Deviation of Unit Weight = 0.004 mm.

PRINCIPAL POINT LOCATION FOR TERRAIN CAMERA, LENS NO. 202

(Emulsion Up)

Direction of Flight

Figure 1.

Master Fiducial Coordinate List for Terrain Camera

(Data provided by Fairchild Company)

Note: All coordinates in millimeters. Refer to Figure 1.

$$A_x = 60.236$$

$$A'_x = 60.262$$

$$A_y = 0.000$$

$$A'_y = -30.778$$

$$B_x = -60.324$$

$$B'_x = -60.384$$

$$B_y = 0.000$$

$$B'_y = 30.843$$

$$C_x = 0.003$$

$$C'_x = 30.887$$

$$C_y = 60.450$$

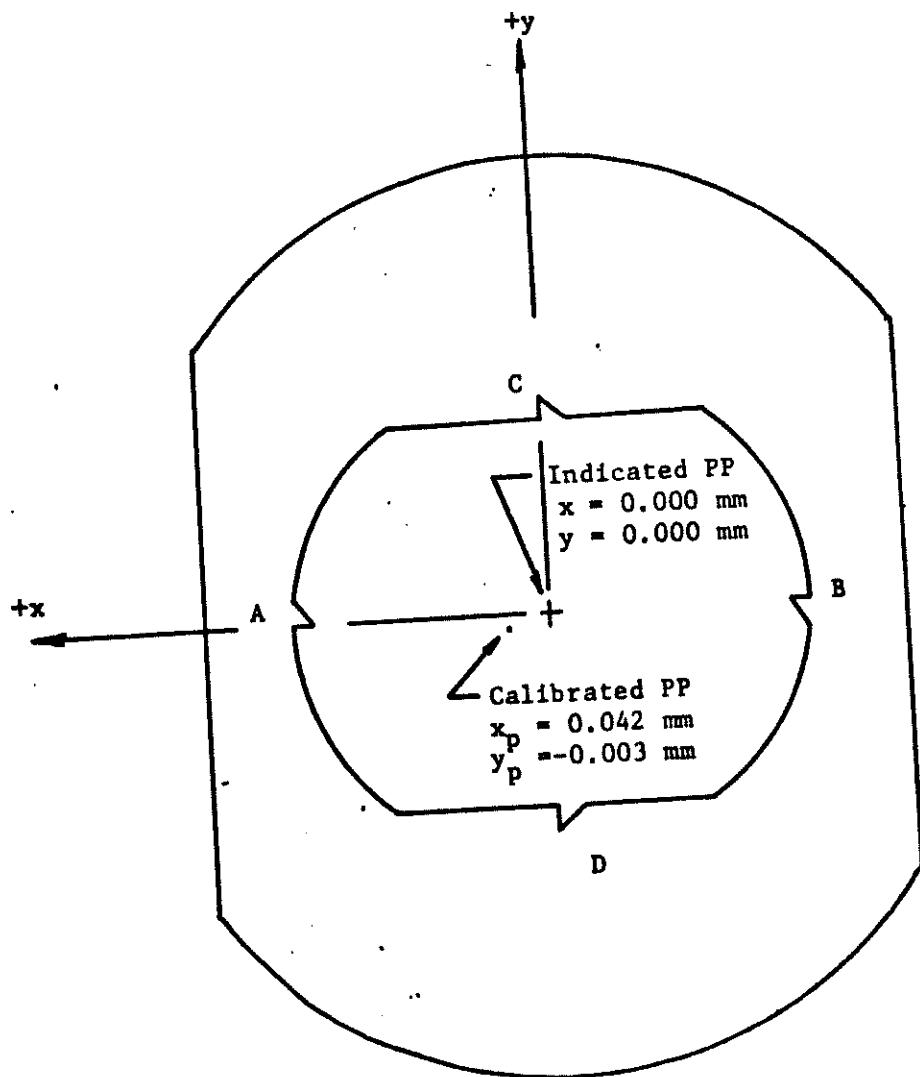
$$C'_y = 60.526$$

$$D_x = -0.003$$

$$D'_x = -30.781$$

$$D_y = -60.305$$

$$D'_y = -60.311$$

PRINCIPAL POINT LOCATION FOR STELLAR CAMERA, LENS NO. 104

(Emulsion up)

Direction of Flight

Figure 2.

EXPLANATORY NOTES - Calibration of Unit SN-003

1. General

All mensuration by Autometric utilized the original negative film stellar exposures.

2. Distortion Function

Radial distortion,  $\Delta r$ , is represented by an odd-power polynomial in  $r$ , the radial distance from the principal point.

$$\Delta r = K_1 r^3 + K_2 r^5 + K_3 r^7$$

The  $x$  and  $y$  components of  $r$  are

$$\Delta x_r = \frac{\Delta r}{r} (x') = (K_1 r^2 + K_2 r^4 + K_3 r^6) (x')$$

$$\Delta y_r = \frac{\Delta r}{r} (y') = (K_1 r^2 + K_2 r^4 + K_3 r^6) (y')$$

where ' $x'$  and ' $y'$ ' are measured image coordinates, relative to principal point origin.

Tangential distortion,  $\Delta t$ , is represented by an even-power polynomial in  $r$ .

$$\Delta t = J_1 r^2 + J_2 r^4$$

The  $x$  and  $y$  components of  $\Delta t$  are

$$\Delta x_t = -\Delta t \sin \phi_0 = -(J_1 r^2 + J_2 r^4) \sin \phi_0$$

$$\Delta y_t = \Delta t \cos \phi_0 = (J_1 r^2 + J_2 r^4) \cos \phi_0$$

where  $\phi_0$  is the angle the axis of maximum tangential distortion makes with the  $x$  axis.

$x'$  and  $y'$  image coordinates can be corrected for radial and tangential lens distortion by the functions

$$x = (1 + K_1 r^2 + K_2 r^4 + K_3 r^6)x' - (J_1 r^2 + J_2 r^4)\sin \phi_0$$

$$y = (1 + K_1 r^2 + K_2 r^4 + K_3 r^6)y' + (J_1 r^2 + J_2 r^4)\cos \phi_0$$

where  $x$  and  $y$  are corrected image coordinates, and  $K_1$ ,  $K_2$ ,  $K_3$ ,  $J_1$ ,  $J_2$ ,  $\phi_0$  are the distortion parameters given by the calibration.

Radial distortion curves for terrain camera lens 202 are presented in Figure 3. The figure gives the EFL radial distortion curve from the stellar calibration and compares the corresponding CFL curve with the Fairchild CFL curve, as determined by laboratory methods. Both CFL curves represent radial distortion characteristics under vacuum conditions, balanced for equal positive and negative distortion values.

Studies made by Fairchild indicate that a negligible change in distortion occurs when the camera is operated under vacuum conditions, rather than the atmospheric conditions under which the stellar calibration was performed. As a result of the supporting Fairchild data given below, no adjustment is made for the change in operating medium from 5000 feet altitude (610 mm. Hg.) to vacuum (0.0001 mm. Hg.).

<u>Field Angle (Degrees)</u>	<u>Distortion Change (Micrometers)</u>
11.25	-0.05
22.50	-0.14
33.75	-0.25
45.00	-1.10

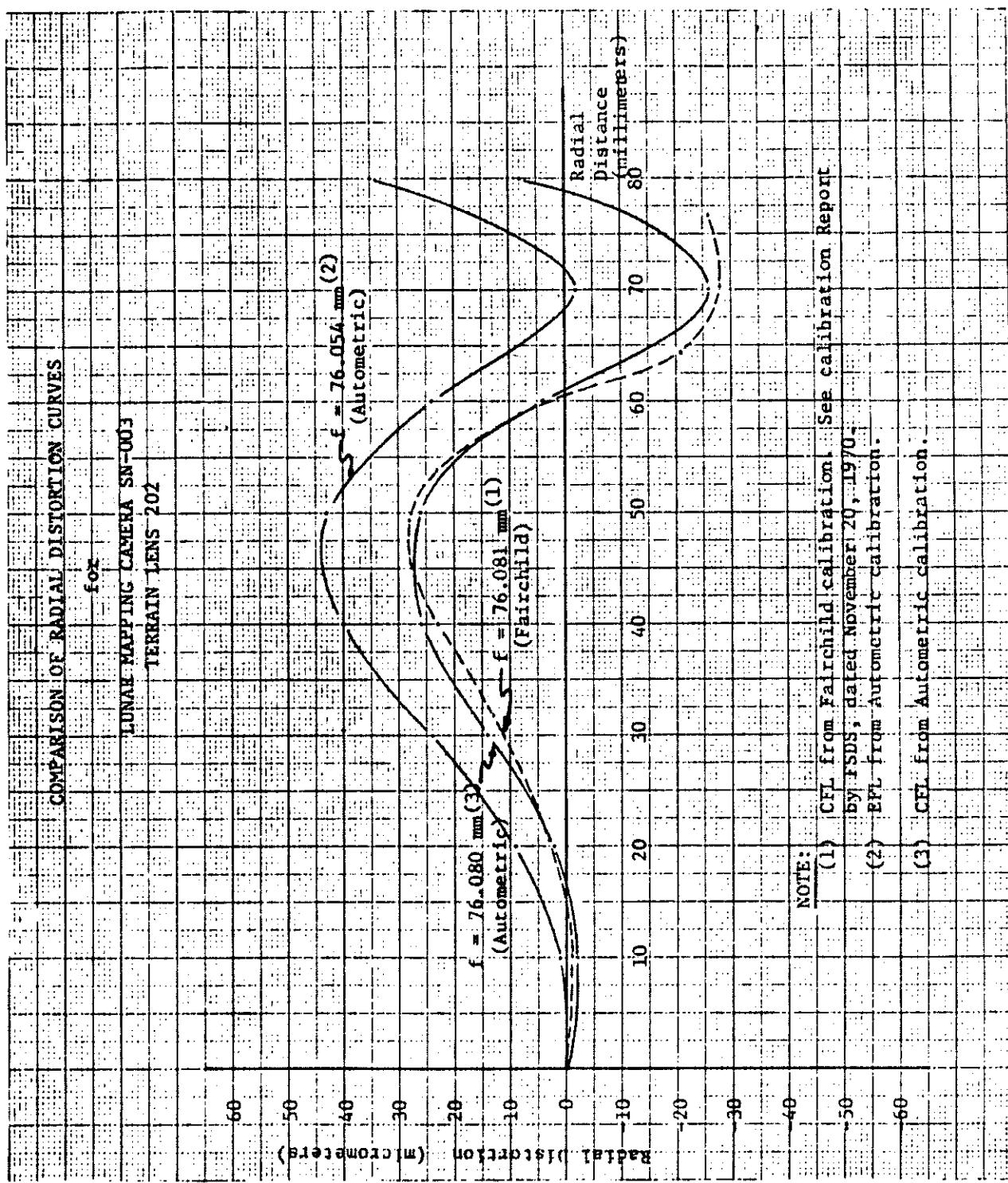


Figure 3.

### 3. Relative Orientation System:

The relative orientation matrix

$$M = \begin{bmatrix} m_{11} & m_{12} & m_{13} \\ m_{21} & m_{22} & m_{23} \\ m_{31} & m_{32} & m_{33} \end{bmatrix}$$

gives the angular orientation of the stellar camera coordinate system with respect to the terrain camera coordinate system. The orientation matrix can be factored into three orthogonal matrices each representing a simple rotation of the stellar camera coordinate system about a particular stellar axis. The sequence of the three rotations must be specified, because different angular orientations result from different sequences. The orientation of  $x_s$ ,  $y_s$ ,  $z_s$  with respect to  $X_T$ ,  $Y_T$ ,  $Z_T$  can be developed as follows.

Consider a stellar camera coordinate system  $x$ ,  $y$ ,  $z$  initially coincident with the terrain camera coordinate system  $X_T$ ,  $Y_T$ ,  $Z_T$  (refer to Figure 4). The three rotations  $\omega$ ,  $\phi$ ,  $\kappa$  are applied to the stellar camera coordinate axes in the given sequence to place the system into its final position,  $x_s$ ,  $y_s$ ,  $z_s$ .

- $\omega$  (roll) - Rotation about the  $x$  axis. Positive  $\omega$  takes the  $+y$  axis toward the  $+z$  axis, resulting in  $x'$ ,  $y'$ ,  $z'$  in Figure 4.
- $\phi$  (pitch) - Rotation about the  $y'$  axis. Positive  $\phi$  takes the  $+z'$  axis toward the  $+x'$  axis, resulting in  $x''$ ,  $y''$ ,  $z''$  in Figure 4.
- $\kappa$  (yaw) - Rotation about the  $z''$  axis. Positive  $\kappa$  takes the  $+x''$  axis toward the  $+y''$  axis, resulting in the final position of the stellar camera coordinate system  $x_s$ ,  $y_s$ ,  $z_s$  in Figure 4.

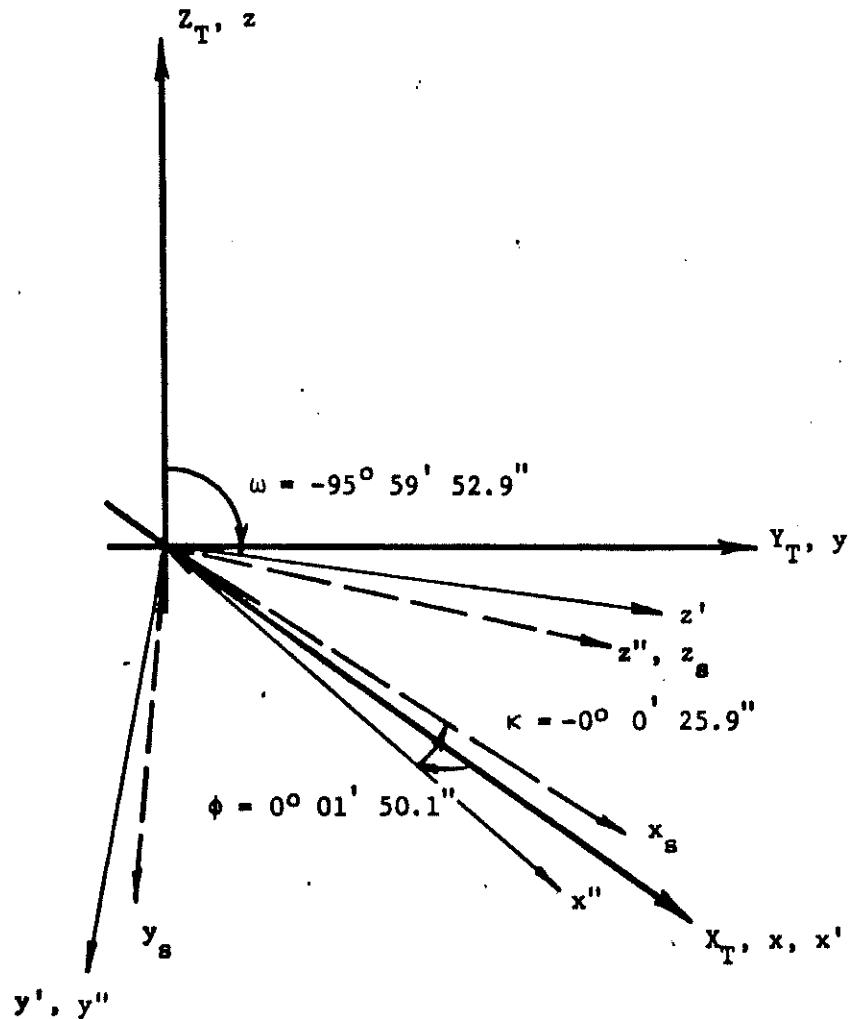


Figure 4. Orientation of Stellar Camera Coordinate System  
With Respect to Terrain Camera Coordinate System.

This Final Report was prepared for Fairchild Space and Defense Systems by Raytheon Company, Autometric Operation, under Contract N-0234.

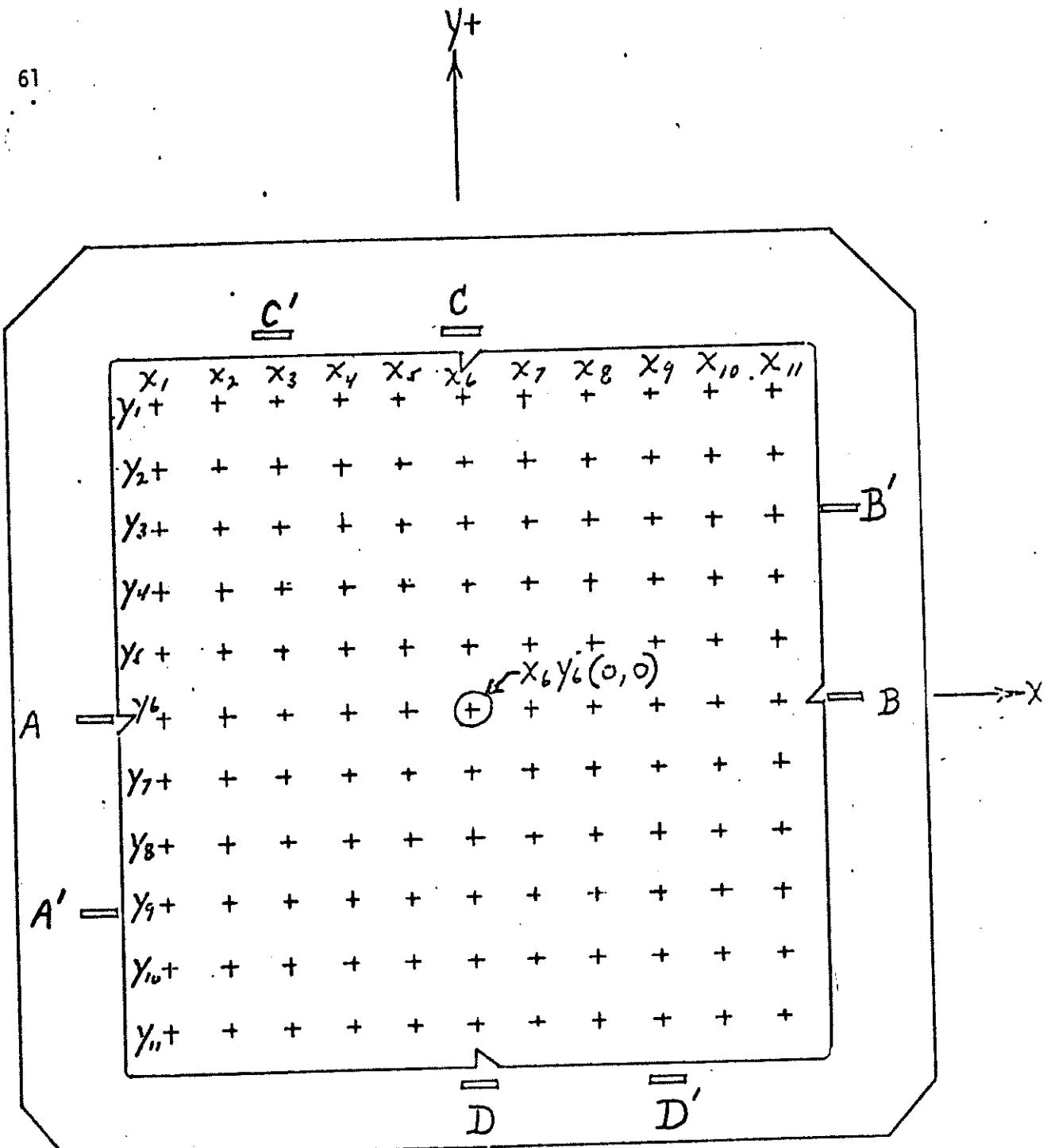
*Ronald K. Brewer*

Ronald K. Brewer  
Senior Scientist/Photogrammetry  
Program Manager

STELLAR CALIBRATION REPORT SUPPLEMENT

Calibrated Ikogon (Mapping)  
and  
Ikotar (Stellar)  
Reseau Coordinate Intersections

IKOGON B GRID PLATE L1Y0U1



{EMULSION UP}

← LINE OF FLIGHT →

**IKOGON B MASTER GRID PLATE**  
**CALIBRATION DATA**  
**( COORDINATES IN MILLIMETERS )**

CROSS	COORDINATES		CROSS	COORDINATES	
	X-	Y-		X-	Y-
X1, Y1	-50.0014	+50.0019	X1, Y4	-50.0014	+20.0015
X2, Y1	-39.9997	+50.0019	X2, Y4	-39.9997	+20.0015
X3, Y1	-29.9977	+50.0019	X3, Y4	-29.9977	+20.0015
X4, Y1	-19.9974	+50.0019	X4, Y4	-19.9974	+20.0015
X5, Y1	- 9.9990	+50.0019	X5, Y4	- 9.9990	+20.0015
X6, Y1	0.0000	+50.0019	X6, Y4	0.0000	+20.0015
X7, Y1	+ 9.9992	+50.0019	X7, Y4	+ 9.9992	+20.0015
X8, Y1	+19.9995	+50.0019	X8, Y4	+19.9995	+20.0015
X9, Y1	+29.9994	+50.0019	X9, Y4	+29.9994	+20.0015
X10, Y1	+40.0019	+50.0019	X10, Y4	+40.0019	+20.0015
X11, Y1	+50.0029	+50.0019	X11, Y4	+50.0029	+20.0015
X1, Y2	-50.0014	+40.0022	X1, Y5	-50.0014	+10.0013
X2, Y2	-39.9997	+40.0022	X2, Y5	-39.9997	+10.0013
X3, Y2	-29.9977	+40.0022	X3, Y5	-29.9977	+10.0013
X4, Y2	-19.9974	+40.0022	X4, Y5	-19.9974	+10.0013
X5, Y2	- 9.9990	+40.0022	X5, Y5	- 9.9990	+10.0013
X6, Y2	0.0000	+40.0022	X6, Y5	0.0000	+10.0013
X7, Y2	+ 9.9992	+40.0022	X7, Y5	+ 9.9992	+10.0013
X8, Y2	+19.9995	+40.0022	X8, Y5	+19.9995	+10.0013
X9, Y2	+29.9994	+40.0022	X9, Y5	+29.9994	+10.0013
X10, Y2	+40.0019	+40.0022	X10, Y5	+40.0019	+10.0013
X11, Y2	+50.0029	+40.0022	X11, Y5	+50.0029	+10.0013
X1, Y3	-50.0014	+30.0030	X1, Y6	-50.0014	0.0000
X2, Y3	-39.9997	+30.0030	X2, Y6	-39.9997	0.0000
X3, Y3	-29.9977	+30.0030	X3, Y6	-29.9977	0.0000
X4, Y3	-19.9974	+30.0030	X4, Y6	-19.9974	0.0000
X5, Y3	- 9.9990	+30.0030	X5, Y6	- 9.9990	0.0000
X6, Y3	0.0000	+30.0030	X6, Y6	0.0000	0.0000
X7, Y3	+ 9.9992	+30.0030	X7, Y6	+ 9.9992	0.0000
X8, Y3	+19.9995	+30.0030	X8, Y6	+19.9995	0.0000
X9, Y3	+29.9994	+30.0030	X9, Y6	+29.9994	0.0000
X10, Y3	+40.0019	+30.0030	X10, Y6	+40.0019	0.0000
X11, Y3	+50.0029	+30.0030	X11, Y6	+50.0029	0.0000

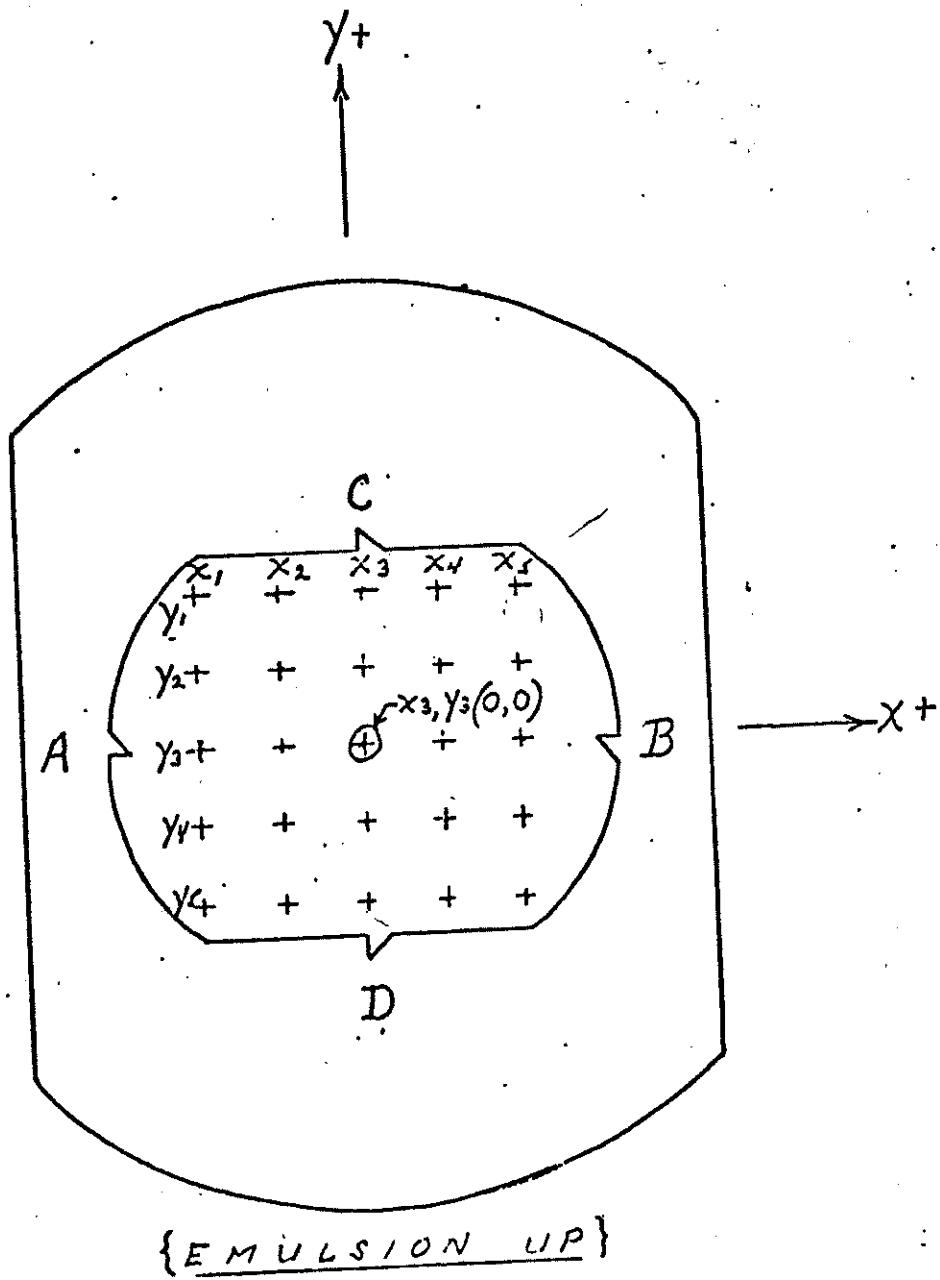
63 IKOGON B MASTER GRID PLATE  
CALIBRATION DATA ----(COORDINATES IN MILLIMETERS)

CROSS	COORDINATES		CROSS	COORDINATES	
	X-	Y-		X-	Y-
X1, Y7	-50.0014	-10.0000	X1, Y10	-50.0014	-39.9974
X2, Y7	-39.9997	-10.0000	X2, Y10	-39.9997	-39.9974
X3, Y7	-29.9977	-10.0000	X3, Y10	-29.9977	-39.9974
X4, Y7	-19.9974	-10.0000	X4, Y10	-19.9974	-39.9974
X5, Y7	- 9.9990	-10.0000	X5, Y10	- 9.9990	-39.9974
X6, Y7	0.0000	-10.0000	X6, Y10	0.0000	-39.9974
X7, Y7	+ 9.9992	-10.0000	X7, Y10	+ 9.9992	-39.9974
X8, Y7	+19.9995	-10.0000	X8, Y10	+19.9995	-39.9974
X9, Y7	+29.9994	-10.0000	X9, Y10	+29.9994	-39.9974
X10, Y7	+40.0019	-10.0000	X10, Y10	+40.0019	-39.9974
X11, Y7	+50.0029	-10.0000	X11, Y10	+50.0029	-39.9974
X1, Y8	-50.0014	-19.9992	X1, Y11	-50.0014	-49.9969
X2, Y8	-39.9997	-19.9992	X2, Y11	-39.9997	-49.9969
X3, Y8	-29.9977	-19.9992	X3, Y11	-29.9977	-49.9969
X4, Y8	-19.9974	-19.9992	X4, Y11	-19.9974	-49.9969
X5, Y8	- 9.9990	-19.9992	X5, Y11	- 9.9990	-49.9969
X6, Y8	0.0000	-19.9992	X6, Y11	0.0000	-49.9969
X7, Y8	+ 9.9992	-19.9992	X7, Y11	+ 9.9992	-49.9969
X8, Y8	+19.9995	-19.9992	X8, Y11	+19.9995	-49.9969
X9, Y8	+29.9994	-19.9992	X9, Y11	+29.9994	-49.9969
X10, Y8	+40.0019	-19.9992	X10, Y11	+40.0019	-49.9969
X11, Y8	+50.0029	-19.9992	X11, Y11	+50.0029	-49.9969
X1, Y9	-50.0014	-29.9987			
X2, Y9	-39.9997	-29.9987			
X3, Y9	-29.9977	-29.9987			
X4, Y9	-19.9974	-29.9987			
X5, Y9	- 9.9990	-29.9987			
X6, Y9	0.0000	-29.9987			
X7, Y9	+ 9.9992	-29.9987			
X8, Y9	+19.9995	-29.9987			
X9, Y9	+29.9994	-29.9987			
X10, Y9	+40.0019	-29.9987			
X11, Y9	+50.0029	-29.9987			

CERTIFICATIONSigned F. Miller

IKOTAIR B GRID PLATE LAYOUT

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{EMULSION UP}

LINE OF FLIGHT

**IKOTAR B MASTER GRID PLATE  
CALIBRATION DATA  
(COORDINATES IN MILLIMETERS)**

CROSS	COORDINATES		CROSS	COORDINATES	
	X-	Y-		X-	Y-
X1, Y1	-10.0015	+10.0003	X1, Y4	-10.0015	-5.0002
X2, Y1	-5.0004	+10.0005	X2, Y4	-5.0008	-5.0005
X3, Y1	-0.0003	+10.0000	X3, Y4	-0.0005	-5.0004
X4, Y1	+4.9996	+10.0000	X4, Y4	+4.9995	-5.0000
X5, Y1	+9.9996	+10.0000	X5, Y4	+9.9996	-5.0002
X1, Y2	-10.0015	+5.0011	X1, Y5	-10.0015	-10.0015
X2, Y2	-5.0004	+5.0001	X2, Y5	-5.0010	-10.0015
X3, Y2	0.0000	+5.0000	X3, Y5	-0.0005	-10.0015
X4, Y2	+4.9996	+5.0006	X4, Y5	+4.9992	-10.0015
X5, Y2	+9.9992	+5.0004	X5, Y5	+9.9992	-10.0015
X1, Y3	-10.0015	+0.0006			
X2, Y3	-5.0009	+0.0002			
X3, Y3	0.0000	+0.0000			
X4, Y3	+4.9993	+0.0006			
X5, Y3	+9.9992	+0.0004			

FAIRCHILD SPACE AND DEFENSE SYSTEMS  
A DIVISION OF FAIRCHILD CAMERA AND INSTRUMENT CORPORATION

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Lens Type IKOGON "B"

Date 12 Oct 70

Serial Number 202

Relative Illumination

<u>Axis</u>	<u>5°</u>	<u>10°</u>	<u>15°</u>	<u>20°</u>	<u>25°</u>	<u>30°</u>	<u>35°</u>	<u>40°</u>	<u>45°</u>
98%	98%	100%	99%	98%	96%	91%	78%	66%	45%

$$\leftarrow = 84.9\%$$

Lens Transmittance = 43.9%

$$\text{Lens T#} = \frac{4.5}{\sqrt{.439}} = \frac{4.5}{663} = 6.787$$

$$\text{Awat} = \frac{6.787}{\sqrt{.849}} = \frac{6.787}{.922}$$

$$\text{Awat} = 7.36$$

Relative Aperture

$$\frac{\text{EFL}}{\text{EFFECTIVE APERTURE}} = \frac{2.0-35''}{.650''} = 4.35$$

RESOLUTION Date: Nov 20, 1970

3.1 Average Resolving Power - Contrast 1000:1

FILM TYPE: 3404

	<u>Axis</u>	<u>7.5°</u>	<u>15°</u>	<u>22.5°</u>	<u>27.5°</u>	<u>30°</u>	<u>32.5°</u>	<u>35°</u>	<u>37.5°</u>	<u>40°</u>	<u>42.5°</u>	<u>45°</u>
Angular	144	150	126	124	105	126	134	134	129	106	55	61
Total	162	148	144	162	162	162	162	135	162	158	119	102

3.2 AWAR = 135.5 L/MM

**FAIRCHILD SPACE AND DEFENSE SYSTEMS**  
**A Division of Fairchild Camera and Instrument Corporation**  
**El Segundo, California**

67

**LENS TYPE** 140TAR "B"

DATE 6 Oct. 1970

SERIAL NUMBER 104

## II. RESOLVING POWER

FILM 3400

PROCESSING Mx641 1 min @ 65° DIAG. #1 AWAR 84.4

TARGET CONTRAST 1000:1      DIAG. #2 AWAR 94.5

DIAG. #1	0°	2½°	5°	7½°	10°						
RADIAL	90	50	59	84	85						
TANGENTIAL	90	95	89	79	75						

DIAG. #2	0°	25°	50°	75°	100°						
RADIAL	90	73	55	84	71						
TANGENTIAL	50	55	35	73	73						

~~FAIRCHILD SPACE AND DEFENSE SYSTEMS~~

A DIVISION OF FAIRCHILD CAMERA & INSTRUMENT CORPORATION

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EL SEGUNDO, CALIFORNIA

Lens Type IKOTAR "B"

Date 6 Oct. 1970

Serial Number 104

Relative Illumination

$\frac{10^{\circ}}{90}$	$\frac{75^{\circ}}{96}$	$\frac{5^{\circ}}{99}$	$\frac{25^{\circ}}{100}$	<u>Axis</u>	$\frac{25^{\circ}}{100}$	$\frac{5^{\circ}}{97}$	$\frac{75^{\circ}}{94}$	$\frac{10^{\circ}}{88}$
-------------------------	-------------------------	------------------------	--------------------------	-------------	--------------------------	------------------------	-------------------------	-------------------------

Transmission

93%

Relative Aperture

1/2.75

## 6.2 CALIBRATION DATA FOR THE OPTICAL BAR PANORAMIC CAMERA

The ITEK calibration reports include relative illumination, spectral transmittance, veiling glare, calibrated focal length and system resolution.

Test Procedure No. TP125

No. of Pages 8

70

## TEST PROCEDURE

FOR

PROJECT 9446  
RELATIVE ILLUMINATION  
MEASUREMENTS  
PANORAMIC CAMERA  
FOR SCIENTIFIC INSTRUMENT  
MODULE

EXPERIMENT S-163

Itek

ITEK CORPORATION

Lexington 73, Massachusetts

Date 9-15-70

Lot no S/N N-54

	PREPARED	PROJECT APPROVAL	QUALITY ASSURANCE APPROVAL
By	R. SHERLOCK	C. BACKE	R. WESPISER
Signed	<u>R. Sherlock</u>	<u>C. Backe</u>	<u>R. Wespiser</u> GA
Date	<u>9/17/70</u>	<u>9/21/70</u>	<u>9/18/70</u>

CUST./GOV'T. REP. \_\_\_\_\_ Date \_\_\_\_\_  
Reviewed

## 5.3.1.3 Itak Test Data Sheet

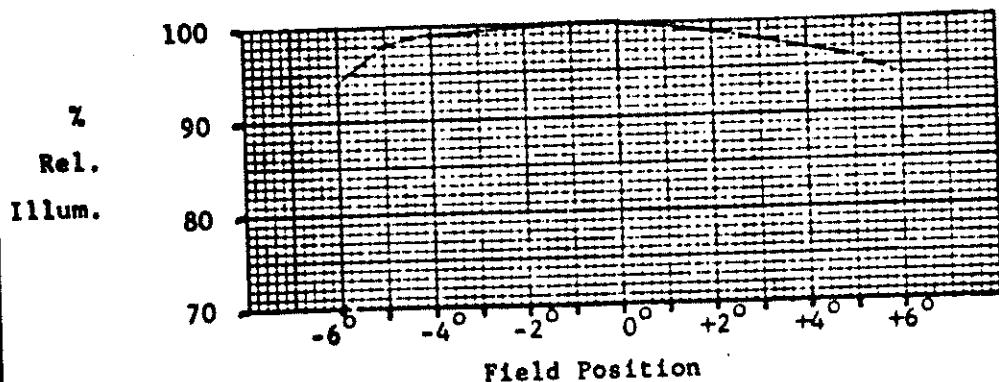
## RELATIVE ILLUMINATION MEASUREMENT

Panoramic Camera Lens, P/N 105150, Serial No. N-54

Field Position	Radiometer Readings vdc	Relative Illumination %
B	-6°	76.0
	-5°	77.5
	-4°	80.0
	-3°	80.5
	-2°	80.8
	-1°	81.0
A	0°	81.0
B	+1°	80.5
	+2°	80.0
	+3°	79.5
	+4°	78.5
	+5°	77.5
	+6°	76.0

Meter Scale for Radiometer Readings: 100

$$\% \text{ Relative Illumination} = \frac{B}{A} \times 100$$

Data Recorded By: Jiter J. Goh Date: 10/12/70QA Monitor: A.C. Delley 10-15-70Project Approval: Craig Bache 10-19-70Test Procedure No. TP125 Page 8

Q199-5 11/65

ITEK CORPORATION  
Lexington 73, Massachusetts

Test Procedure No. TP123No. of Pages 9**TEST PROCEDURE****FOR**

PROJECT 9446  
 SPECTRAL TRANSMITTANCE  
 MEASUREMENTS  
 PANORAMIC CAMERA  
 FOR SCIENTIFIC INSTRUMENT  
 MODULE

EXPERIMENT S-163

ITEK CORPORATION

Lexington 73, Massachusetts

Date 9-15-70

Loc S/W N-54

	PREPARED	PROJECT APPROVAL	QUALITY ASSURANCE APPROVAL
By	R. SHERLOCK	C. BACKE	R. WESPISER
Signed	<u>R. Sherlock</u>	<u>C. Backe</u>	<u>R. Wespiser</u>
Date	<u>9/17/70</u>	<u>9/21/70</u>	<u>9/17/70</u>

CUST./GOV'T. REP. \_\_\_\_\_ Date \_\_\_\_\_

Reviewed \_\_\_\_\_

## 6.3.1.3 Itek Test Data Sheet

## SPECTRAL TRANSMITTANCE MEASUREMENT

Panoramic Camera Lens, P/N 105150, Serial No. 11-54

Wavelength nm	Radiometer Readings				% Transmittance
	A	A <sub>1</sub>	B	B <sub>1</sub>	
400	9.2	2.65	2.0	.13	22.6
420	25.8	6.6	5.4	.55	33.8
440	46.0	12.2	9.5	1.3	51.6
480	100.0	27.0	21.0	3.8	67.0
520	144.9	41.0	32.5	6.6	71.8
560	100.0	44.5	36.0	7.4	73.9
601	145.0	59.5	33.0	6.2	69.0
640	125.0	35.0	21.5	4.5	63.8
680	106.0	25.7	19.0	3.5	56.7
720	86.0	18.0	13.0	2.0	47.4

## Legend:

A = Brightness of the calibrated standard Lambertian source using the radiometer telescope.

A<sub>1</sub> = Brightness of the collimator target as seen from the lens test position using the radiometer telescope.

B = Brightness of the calibrated standard Lambertian source using the radiometer microscope.

B<sub>1</sub> = Brightness of the collimator target aerial image at the image plane (lens in place) using the radiometer microscope.

$$\% \text{ Transmittance} = [(B_1 / B) \div (A_1 / A)] \times 100.$$

Data Recorded By: Itek/Stack Date: 4/12/70

QA Monitor: J.C. McHugh Date: 4-15-70

Project Approval: Conrad Date: 4/18/70

Test Procedure No. TP123

Page 8

## 6.3.1.3 Itek Test Data Sheet

## T STOP CALCULATION

Panoramic Camera Lens, P/N 105150, Serial No. N-54

Wavelength nm	% Transmittance	T STOP
400	22.6	7.4
420	39.8	5.5
440	51.6	4.9
480	67.0	4.3
520	71.5	4.1
560	73.7	4.1
601	69.0	4.2
640	63.8	4.4
680	56.7	4.6
720	49.4	5.0

## Legend:

$$T \text{ STOP} = \frac{f/\text{number}}{\sqrt{t}} \quad \text{where } f/\text{number} \text{ is 3.5 and}$$

t is transmittance.

Data Recorded By:

QA Monitor:

Project Approval:

Eugen J. StachA.C. DeKeyCarl BrackDate: OCT 12 197010-15-7010-19-70Test Procedure No. TP123

Q199-5 11/65

Page 9ITEK CORPORATION  
Lexington 73, Massachusetts

Test Procedure No. TP121

No. of Pages 8**TEST PROCEDURE****FOR**

PROJECT 9446

VEILING GLARE MEASUREMENTS  
PANORAMIC CAMERA  
FOR SCIENTIFIC INSTRUMENT  
MODULE

EXPERIMENT S-163

Itek

ITEK CORPORATION

Lexington 73, Massachusetts

Date 9-15-70

Lens S/N N-54

	PREPARED	PROJECT APPROVAL	QUALITY ASSURANCE APPROVAL
By	R. SHERLOCK	C. BACKE	R. WESPISER
Signed	<i>R. Sherlock</i>	<i>C. Backe</i>	<i>R. Wespiser</i>
Date	<u>9/17/70</u>	<u>9/21/70</u>	<u>9/18/70</u>

CUST./GOV'T. REP.

Date

Reviewed

## 6.3.1.3 Itek Test Data Sheet

## VEILING GLARE MEASUREMENT

Panoramic Camera Lens, P/N 105150, Serial No. H-54

Step No.	Step Wedge							Black Dot %	Veiling Glare %
	1	2	3	4	5	6	7		
* Calibrated Value									
	1.22	1.07	.92	.78	.65	.50	.36		
* Measured Test Values									
-6°	.24	.34	.51	.20	1.14	1.44	1.79	.55	12.6
-4°	.22	.35	.43	.65	.94	1.26	1.62	.44	12.3
-2°	.60	.13	1.32	1.62	1.88	2.04	2.24	1.34	12.7
0°	.24	.32	.50	.80	1.13	1.44	1.74	.54	13.2
+2°	.26	.36	.55	.84	1.11	1.40	1.77	.82	16.2
+4°	.25	.37	.56	.86	1.14	1.45	1.83	.84	16.2
+6°	.22	.30	.46	.71	1.05	1.36	1.70	.64	15.1

Legend: \* Density values are logarithmic as read on the Macbeth Densitometer.

% Veiling Glare is computed from the black spot density plot (attached) for each field position.

Data Recorded By:

John Hall Date: 10/12/70

QA Monitor:

JKH/SJ Date: 10/16/70

Project Approval:

Carl Bach Date: 10/19/70Test Procedure No. TP121.

Q199-5 11/65

Page 8ITEK CORPORATION  
Lexington 73, Massachusetts

Test Procedure No. TP127No. of Pages 5**TEST PROCEDURE****FOR**

PROJECT 9446

C.F.L./MEASUREMENTS  
PANORAMIC CAMERAFOR SCIENTIFIC INSTRUMENT  
MODULE  
EXPERIMENT S-163**Itek**

ITEK CORPORATION

Lexington 73, Massachusetts

Date 9-15-70

	PREPARED	PROJECT APPROVAL	QUALITY ASSURANCE APPROVAL
By	R. SHERLOCK	C. BACKE	R. WESPISER
Signed	<i>R. SHERLOCK</i>	<i>C. Backe</i>	<i>R. Wespiser</i>
Date	<u>9/17/70</u>	<u>9/21/70</u>	<u>✓ 9/21/70</u>

CUST./GOV'T. REP.

Reviewed

Date

### 6.3.1.3 Test Data Summary

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Lens - N-54

#### C.F.L. Calibration Summary

Filter	Field Position (degrees)	Mean C.F.L. (inches)	STD Deviation of of mean (C.F.L.)
23A	+6	24.0057	0.0004
	+4	24.0058	0.0003
	+2	24.0015	0.0006
	0 *	24.0052	0.0008
	-2	24.0073	0.0006
	-4	24.0057	0.0001
	-6	24.0050	0.0003
12	+6	24.0063	0.0008
	+4	24.0080	0.0003
	+2	24.0086	0.0009
	0 *	24.0081	0.0004
	-2	24.0088	0.0006
	-4	24.0089	0.0004
	-6	24.0080	0.0001
8	+6	24.0076	0.0004
	+4	24.0096	0.0007
	+2	24.0106	0.0005
	0 *	24.0080	0.0007
	-2	24.0059	0.0003
	-4	24.0081	0.0015
	-6	24.0064	0.0009
2A	+6	24.0044	0.0003
	+4	24.0075	0.0006
	+2	24.0085	0.0008
	0 *	24.0062	0.0007
	-2	24.0047	0.0013
	-4	24.0055	0.0006
	-6	24.0066	0.0005
no filter	+6	24.0037	0.0003
	+4	24.0050	0.0004
	+2	24.0065	0.0006
	0 *	24.0045	0.0005
	-2	24.0032	0.0004
	-4	24.0048	0.0008
	-6	24.0037	0.0006

Avg: 24.0045 in = 609.717 mm

\* Average of all field positions

Quality Assurance Review

A. C. McKey QA <sub>32</sub>

Test Procedure No. TP-70-9446-42-4B

No. of Pages 47

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3/N003

1-27-71

# ITEK APPROVED

## TEST PROCEDURE

### FOR

PHOTOGRAPHIC BASELINE (PBL)

(DYNAMIC TESTING WITH FORWARD MOTION COMPENSATION)

ACCEPTANCE OF THE PANORAMIC CAMERA

EXPERIMENT S-163

AESU 23A Filter.



ITEK CORPORATION

Lexington 73, Massachusetts

Date 9/23/70

	PREPARED	PROJECT APPROVAL	QUALITY ASSURANCE APPROVAL
By	D. D. Auer	W. B. Lyle	W. Zedley
Signed	C. C. Tandy	J. F. Lyle	W. Zedley
Date	7/24/70	7/24/70	7/24/70

CUST./GOV'T. REP. \_\_\_\_\_ Date \_\_\_\_\_  
Reviewed \_\_\_\_\_

Q199-1 11/65

NO FILTER  
FLIGHT CONFIG.

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## DYNAMIC TESTING WITH FMC w/o FILTER

**DATA SHEET**

RUN 95/G

TP-70-9446-42-4

PAGE 37-3

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## DATA SHEET

Frieder Poetsch et al.

卷之三

*Hippoboscidae* G.F.  
Genus *Hippoboscidae*

out of film  
- yet they wrote  
- 6 people  
using + 4°  
using - 4°

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TP-70-9446-42-4-B

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## 7.0 PHOTOGRAPHIC MISSION OPERATIONS

Mission operation contingencies required rescheduling of some photography during the latter part of the mission; however, the photography accomplished generally covered the areas delineated in premission planning. Operation of the laser altimeter was curtailed because of a malfunction, thus all scheduled operations from Rev 50 on were accomplished on daylight side only. A complete girth of the moon with the laser was recorded on Rev 15/16 with sporadic data being recorded the remaining time in lunar orbit. No estimate of total operational laser data recorded can be made until postflight analysis is completed.

On the accompanying figures the times shown in the "accomplished" column are flight plan update times made during the mission. The longitudes opposite these times were determined from mission trajectories and are included merely as a reference as they are only approximate values. The time and position for each exposure will be listed in the orbital support data (APE) that is to be published.

Major deviations from the planned mission are listed below.

### Metric Camera

Rev 35 - North oblique strip was flown with a spacecraft attitude that introduced yaw in the camera orientation and are not true north obliques as planned.

Rev 50 - Camera turned off early as spacecraft attitude had drifted beyond desired limits.

Rev 62 - Camera turned on late because of other activities. This is a reschedule of photography missed on Rev 58/59. Lightside pass only.

Rev 63 - Reduced to lightside pass only.

Rev 70 - Reschedule of Rev 69. Lightside pass only.

Rev 72 - Lightside pass only.

Panoramic Camera - Telemetry readouts received at MSC during the mission indicated the FMC was giving spurious readings at an <sup>err</sup> iratic interval. Estimates were that approximately 80% of the time the unit was functioning correctly.

Rev 38 - Last part of first strip on this rev was switched to mono; however, time of this change not known.

Rev 58 - Not accomplished because of extended rest period.

Rev 60 - Partial reschedule of Rev 58. Second test cycle taken over LM impact area.

Rev 61 Additional photos over LM impact area to provide stereo photography with the Rev 60 photos.

Rev 63 - Partial reschedule of Rev 58.

## SCHEDULED

REV	START		STOP		REMARKS	ACCOMPLISHED		STOP	REMARKS
	GET(H:M:S)	LONG(DEC)	GET(H:M:S)	LONG(DEC)		GET(H:M:S)	LONG(DEC)		
3/4	84:42:35	179.6W	84:54:15	142.5E		4	84:42:23	179.1W	84:54:14
15	105:52:20	144.5E	106:17:20	74.0E		15	105:52:58	144.4E	106:17:57
15/16	106:56:15	49.8W	108:54:40	50.8W	Full Rev	15/16	106:56:51	50.0W	108:55:19
22/23	119:33:35	162.7E	121:32:02	161.7E	Full Rev	22/23	119:34:53	161.3E	121:33:02
23	121:37:04	143.7E	122:31:28	18.8W	Forward	23	121:39:34	141.5E	122:32:43
27	129:26:27	157.7E	130:26:00	22.8W		27	129:26:52	157.2E	130:26:30
33/34	141:16:34	151:6E	143:15:02	150.6E	Full Rev	33/34	141:17:26	150.8E	143:14:00
34	143:15:02	150.6E	144:09:00	11.8W	BKWD Ob.		144:09:30	144:09:30	9.5W
35	145:13:30	149.6E	146:12:55	30.8W	N Ob.	35	145:14:16	149.6E	146:13:56
37/38	150:08:58	32.9W	152:08:20	33.9W	Full Rev	38	151:09:22	146.6E	152:09:32
44	162:59:44	140.6E	163:59:10	39.9W		44	162:59:30	140.5E	163:59:11
50	174:50:06	134.6E	175:49:22	45.9W		50	174:50:04	134.4E	175:32:32
58/59	190:37:25	126.6E	192:35:51	125.6E	Full Rev	62	198:44:40	83.4E	199:31:56
62/63	199:30:22	57.9W	201:28:11	58.9W	Full Rev	63	200:30:56	120.0E	201:30:21
69	212:19:43	115.6E	214:18:10	114.6E	Full Rev	70	214:20:03	114.5E	215:19:49
71	216:16:37	113.6E	217:15:47	66.9W	South Ob	71	216:18:30	113.4E	217:18:15
71/72	217:20:00	82.9W	219:14:14	67.9W		72	218:16:59	112.3E	219:16:44

FIG. 7-1 Apollo 15 Mapping Camera Photography

PANORAMIC CAMERA  
PHOTOGRAPHY

SCHEDULED		ACCOMPLISHED								
REV		START	STOP	REV	START	STOP				
	GET(H:M:S)	LONG(DEC)	GET(H:M:S)	LONG(DEC)	MODE	GET(H:M:S)	LONG(DEC)	GET(H:M:S)	LONG(DEC)	MODE
3/4	84:42:35	179.6W (E TERM)	84:45:55	170.5E	Mono	3/4	84:42:23	179.1W	84:45:43	169.6E
	84:45:55	170.5E	84:54:15	142.5E	Stereo	15	105:52:58	144:4E	106:08:28	100.1E
15	105:52:20	144:5E	106:07:30	100E	Stereo		106:08:28	100.1E	106:14:58	81.8E
	106:07:50	100E	106:14:20	82E	Mono		106:14:58	81.8E	106:17:57	74.0E
16	106:14:20	82E	106:17:20	74E	Stereo	16	108:15:27	75.9E	108:43:15	11.9W
	108:14:40	76E	108:39:15	1.8W	Stereo		130:18:05	5.9E	130:19:16	1.1E
	108:39:15	1.8W	108:42:33	11.8W (W Term)	Mono	27				
27	130:16:45	5.7E	130:18:00	1.7E	Stereo	33	141:46:11	67.5E	142:01:31	23.2E
	141:45:05	67.5E	142:00:20	23E	Stereo	38	151:13:13	133.8E	151:37:01	65.4E
33	151:12:40	134E	151:31:50	77.5E	Stereo	50	175:34:38	4.9E	175:36:56	2.4W
	151:31:50	77.5E	151:36:12	65.5E	Mono		195:04:13	40.7E	195:14:30	12.0E
38	151:56:20	5.7E	151:57:05	1.7E	Stereo	60	195:16:23	6.3E	Test cycle	5 frames
	175:34:07	5.7E	175:36:00	0.3W	Stereo		195:18:23	0.1E	Test cycle	5 frames
58	190:37:25	126.6E (E TERM)	190:40:45	116.6E	Mono	61	197:16:23	0.2E	197:17:43	2.8W
	190:40:45	116.6E	190:55:50	69.E	Stereo		197:17:43	2.8W	197:18:23	4.3W
	190:55:50	69.E	191:01:45	53.E	Mono	63	200:33:29	121.1E	200:49:32	64.8E
	191:01:45	53.E	191:08:20	36.5E	Stereo		200:49:32	64.8E	200:54:27	51.7E
						200:54:27	51.7E	200:59:22	38.0E	Stereo
59	193:05:25	38.5E	193:18:30	0	Stereo	72	218:52:55	9.2E	219:13:24	56.7W
	193:31:45	44.9W	193:35:05	54.9W (W TERM)	Mono		219:13:24	56.7W	219:16:44	70.8W
70	214:42:10	42E	214:50:12	20E	Stereo					
	218:48:30	22E	219:10:54	57.9W	Stereo					
	219:10:54	57.9W	219:14:14	67.9W	Mono					

FIG. 7-2  
Apollo 15 Panoramic Photography

\*\* Rev 70 will be scheduled only if it is determined  
REAL TIME that film is available

NASA — MSC — Coml., Houston, Texas

$\frac{1}{2} \times 100 = 50$

2 ft. Sac/foot

2 ft.

$$\left( 1 - \frac{2}{100} \right) = 0.8$$

So,  $0.8 \times 50 = 40$

So,  $40 \times 100 = 4000$

16. Poured Vol.

100% Sac.

$$100 \times 100 = 10000 \times \left( 1 + \frac{2}{100} \right)^3 = 12544$$

So,  $12544 \times 0.8 = 10035.2$  (approx.)

*Apollo 15, wwp/23/71  
mapping camera*

## 6.0 CAMERA CALIBRATION DATA

Calibration of the Apollo 15 SIM bay cameras was provided by the equipment manufacturers. Copies of the original calibration reports are presented in this section. Although portions of the data may be difficult to read, it seemed more appropriate to include the original reports than to have them retyped. No additional calibration was accomplished by NASA/MSC.

### 6.1 CALIBRATION DATA FOR THE MAPPING CAMERA SUBSYSTEM (MCS)

The stellar calibration report was prepared by Raytheon/Autometric under contract to Fairchild Space and Defense System. In addition to the camera calibrations the report includes calibrated values of the reseaus and the relative orientation angle between the connected lens cones. Fairchild data concerning relative illumination and resolving power of the cameras are included.

TR-71-3404-2R

FINAL

CAMERA CALIBRATION REPORT

CAMERA UNIT SN-003

JULY 1971

Prepared by

Autometric Operation  
Equipment Division  
Raytheon Company

4217 Wheeler Avenue  
Alexandria, Virginia 22304

Prepared for

Fairchild Space and Defense Systems  
300 Robbins Lane  
Syosset, L.I., New York 11791

## FOREWARD

This final stellar calibration report is submitted to NASA Manned Spacecraft Center, Houston, Texas, by Fairchild Space and Defense Systems (FSDS), A Division of Fairchild Camera and Instrument Corporation under NASA Contract No. NAS 9-10457. This report was prepared by Raytheon/Autometric, Alexandria, Virginia, under separate contract to FSDS. It supersedes the preliminary version which was previously submitted in May 1971.

This report summarizes the results of the stellar calibration of Mapping Camera Subsystem (MCS) Serial Number 003. The actual stellar field calibration was performed at the NASA White Sands Test Facility (WSTF), Las Cruces, New Mexico, during the period 25-26 March 1971. The resulting stellar imagery was subsequently reduced by Raytheon/Autometric and computer processed and outputted into the form presented herein. It should be emphasized that the calibration values shown for the individual terrain lens/stellar lens internal geometry parameters and relative orientation between connected lens cones are indicative of the capabilities of MCS 003, as derived by stellar methods. Also included in this report are the calibrated values of the terrain and stellar reseau coordinate intersections.

All calibration values presented herein should be utilized in the subsequent post-flight data reductions of MCS 003 terrain and stellar imagery.



A. D. Beccasio  
Senior Staff Engineer

Approved by: A. G. Hutchins  
A. G. Hutchins  
MCS Program Manager

12 July 1971  
Fairchild Space & Defense Systems  
300 Robbins Lane  
Syosset, New York 11791

## CAMERA CALIBRATION RESULTS

Lunar Mapping Camera SN-003 Stellar Calibration

### Terrain Lens (202) Constants of Internal Geometry

$$EFL = 76.054 \text{ mm.}$$

$$\text{S.D.} = 0.002 \text{ mm.}$$

$$CFL = 76.080 \text{ mm.}$$

$$\text{S.D.} = 0.002 \text{ mm.}$$

Principal Point With Respect to Indicated Principal Point

(Indicated principal point  $x_{ipp} = 0.0 \text{ mm.}, y_{ipp} = 0.0 \text{ mm.})$ ,

$$x_p = -0.006 \text{ mm.}$$

$$\text{S.D.} = 0.001 \text{ mm.}$$

$$y_p = -0.002 \text{ mm.}$$

$$\text{S.D.} = 0.001 \text{ mm.}$$

### Radial Distortion Parameters

$$k_1 = -0.13361854 \times 10^{-5} \quad \text{S.D.} = 0.39272852 \times 10^{-7}$$

$$k_2 = 0.52261757 \times 10^{-9} \quad \text{S.D.} = 0.15933971 \times 10^{-10}$$

$$k_3 = -0.50728336 \times 10^{-13} \quad \text{S.D.} = 0.19335059 \times 10^{-14}$$

### Lens Decentration Distortion Parameters

$$j_1 = -0.54958195 \times 10^{-6} \quad \text{S.D.} = 0.48922291 \times 10^{-6}$$

$$j_2 = -0.46089420 \times 10^{-10} \quad \text{S.D.} = 0.11215963 \times 10^{-9}$$

$$\phi_0 = 2.9659070 \text{ radians} \quad \text{S.D.} = 0.23946572$$

Stellar Lens (104) Constants of Internal Geometry

EFL = 75.713 mm. S.D. = 0.011 mm.  
CFL = 75.735 mm. S.D. = 0.011 mm.

~~Principal Point With Respect to Indicated Principal Point  
(Indicated principal point  $x_{ipp} = 0.0$  mm.,  $y_{ipp} = 0.0$  mm.),~~

$x_p = 0.042$  mm. S.D. = 0.019 mm.  
 $y_p = -0.003$  mm. S.D. = 0.018 mm.

Radial Distortion Parameters

$K_1 = -0.92376425 \times 10^{-5}$  S.D. =  $0.40956407 \times 10^{-5}$   
 $K_2 = 0.53810549 \times 10^{-7}$  S.D. =  $0.33179901 \times 10^{-7}$   
 $K_3 = -0.70032005 \times 10^{-10}$  S.D. =  $0.80049661 \times 10^{-10}$

Lens Decentration Distortion Parameters

$J_1 = 0.81424290 \times 10^{-6}$  S.D. =  $0.10033483 \times 10^{-4}$   
 $J_2 = -0.20686782 \times 10^{-7}$  S.D. =  $0.42974779 \times 10^{-7}$   
 $\phi_0 = 1.2355515$  radians S.D. = 0.75665362

Results of Lock-Angle Calibration

Relative Orientation Matrix Defining a Transformation from the Terrain Camera to the Stellar Camera

0.999999850	-0.000517921	0.000180754
0.000125645	-0.104494099	-0.994525499
0.000533973	0.994525372	-0.104494018

Relative Orientation Angles (Degrees, Minutes, Seconds)

OMEGA =	-95	59	52.859
PHI =	0	1	50.140
KAPPA =	-0	0	25.916

Covariance Matrix

$0.98562 \times 10^{-11}$	$-0.33758 \times 10^{-12}$	$0.88273 \times 10^{-12}$
$-0.33758 \times 10^{-12}$	$0.22323 \times 10^{-10}$	$0.34486 \times 10^{-12}$
$0.88273 \times 10^{-12}$	$0.34486 \times 10^{-12}$	$0.35148 \times 10^{-9}$

Standard Deviation of Orientation Angles (Arc-seconds)

S.D. OMEGA =	0.65
S.D. PHI =	0.98
S.D. KAPPA =	3.87

Statistical Data From Simultaneous Solution

Weighted Sum of Squares = 0.031920

Degrees of Freedom = 1996

Standard Deviation of Unit Weight = 0.004 mm.

PRINCIPAL POINT LOCATION FOR TERRAIN CAMERA, LENS NO. 202

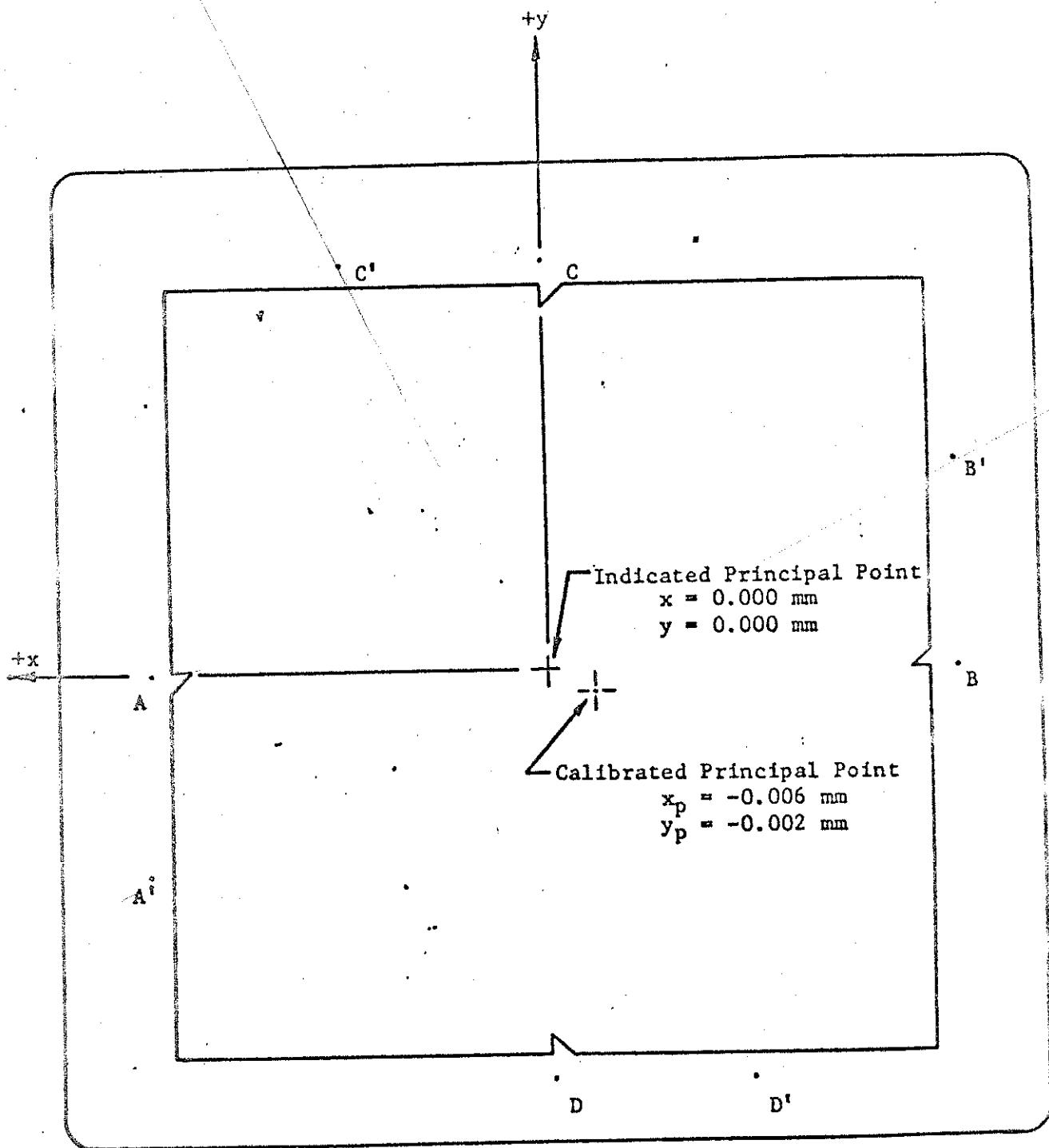


Figure 1.

Master Fiducial Coordinate List for Terrain Camera

(Data provided by Fairchild Company)

Note: All coordinates in millimeters. Refer to Figure 1.

$$A_x = 60.236$$

$$A'_x = 60.262$$

$$A_y = 0.000$$

$$A'_y = -30.778$$

$$B_x = -60.324$$

$$B'_x = -60.384$$

$$B_y = 0.000$$

$$B'_y = 30.843$$

$$C_x = 0.003$$

$$C'_x = 30.887$$

$$C_y = 60.450$$

$$C'_y = 60.526$$

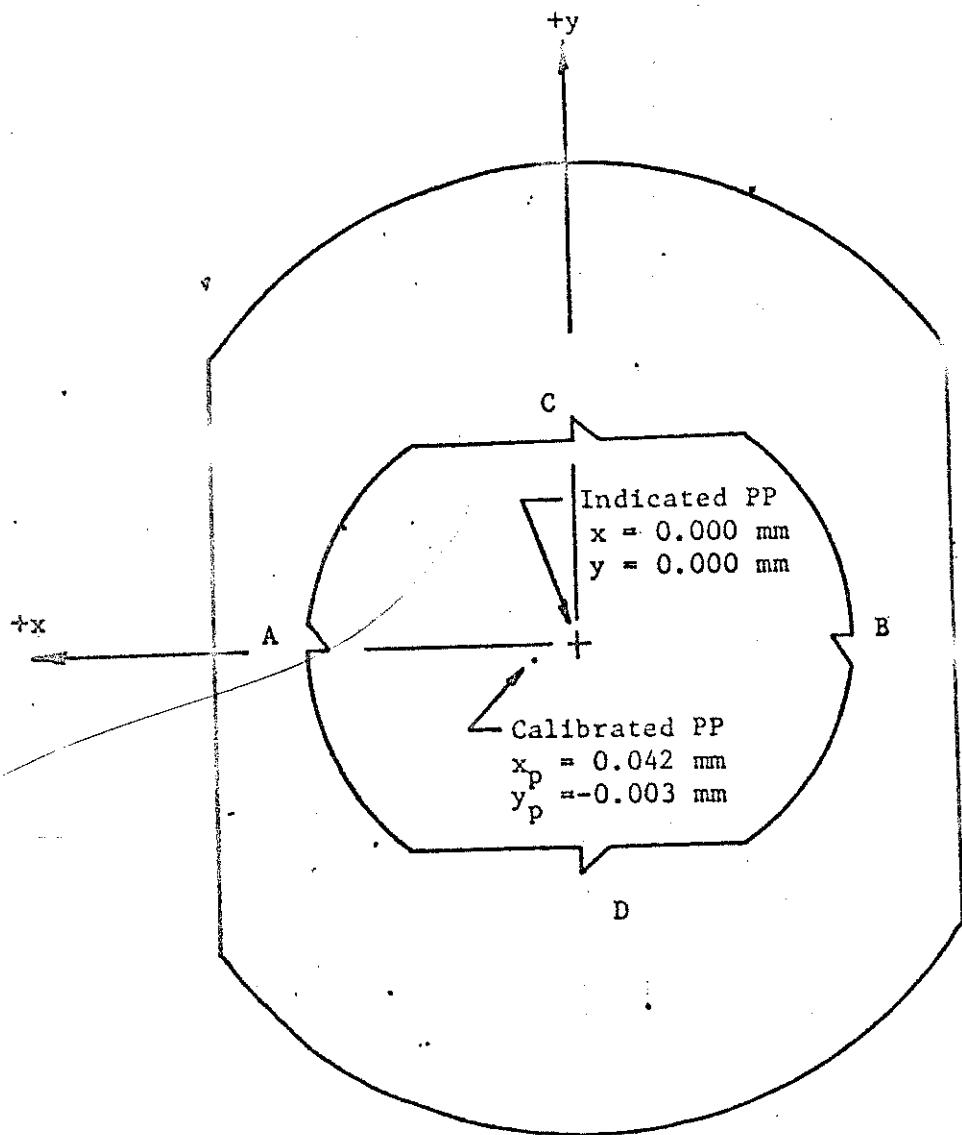
$$D_x = -0.003$$

$$D'_x = -30.781$$

$$D_y = -60.305$$

$$D'_y = -60.311$$

PRINCIPAL POINT LOCATION FOR STELLAR CAMERA, LENS NO. 104



(Emulsion up)

Direction of Flight

Figure 2.

EXPLANATORY NOTES - Calibration of Unit SN-003

1. General

All mensuration by Autometric utilized the original negative film stellar exposures.

2. Distortion Function

Radial distortion,  $\Delta r$ , is represented by an odd-power polynomial in  $r$ , the radial distance from the principal point.

$$\Delta r = K_1 r^3 + K_2 r^5 + K_3 r^7$$

The  $x$  and  $y$  components of  $r$  are

$$\Delta x_r = \frac{\Delta r}{r} (x') = (K_1 r^2 + K_2 r^4 + K_3 r^6) (x')$$

$$\Delta y_r = \frac{\Delta r}{r} (y') = (K_1 r^2 + K_2 r^4 + K_3 r^6) (y')$$

where ' $x'$  and ' $y'$ ' are measured image coordinates, relative to principal point origin.

Tangential distortion,  $\Delta t$ , is represented by an even-power polynomial in  $r$ .

$$\Delta t = J_1 r^2 + J_2 r^4$$

The  $x$  and  $y$  components of  $\Delta t$  are

$$\Delta x_t = -\Delta t \sin \phi_0 = -(J_1 r^2 + J_2 r^4) \sin \phi_0$$

$$\Delta y_t = \Delta t \cos \phi_0 = (J_1 r^2 + J_2 r^4) \cos \phi_0$$

where  $\phi_0$  is the angle the axis of maximum tangential distortion makes with the  $x$  axis.

$x'$  and  $y'$  image coordinates can be corrected for radial and tangential lens distortion by the functions

$$x = (1 + K_1 r^2 + K_2 r^4 + K_3 r^6)x' - (J_1 r^2 + J_2 r^4)\sin \phi_0$$

$$y = (1 + K_1 r^2 + K_2 r^4 + K_3 r^6)y' + (J_1 r^2 + J_2 r^4)\cos \phi_0$$

where  $x$  and  $y$  are corrected image coordinates, and  $K_1$ ,  $K_2$ ,  $K_3$ ,  $J_1$ ,  $J_2$ ,  $\phi_0$  are the distortion parameters given by the calibration.

Radial distortion curves for terrain camera lens 202 are presented in Figure 3. The figure gives the EFL radial distortion curve from the stellar calibration and compares the corresponding CFL curve with the Fairchild CFL curve, as determined by laboratory methods. Both CFL curves represent radial distortion characteristics under vacuum conditions, balanced for equal positive and negative distortion values.

Studies made by Fairchild indicate that a negligible change in distortion occurs when the camera is operated under vacuum conditions, rather than the atmospheric conditions under which the stellar calibration was performed. As a result of the supporting Fairchild data given below, no adjustment is made for the change in operating medium from 5000 feet altitude (610 mm. Hg.) to vacuum (0.0001 mm. Hg.).

<u>Field Angle (Degrees)</u>	<u>Distortion Change (Micrometers)</u>
11.25	-0.05
22.50	-0.14
33.75	-0.25
45.00	-1.10

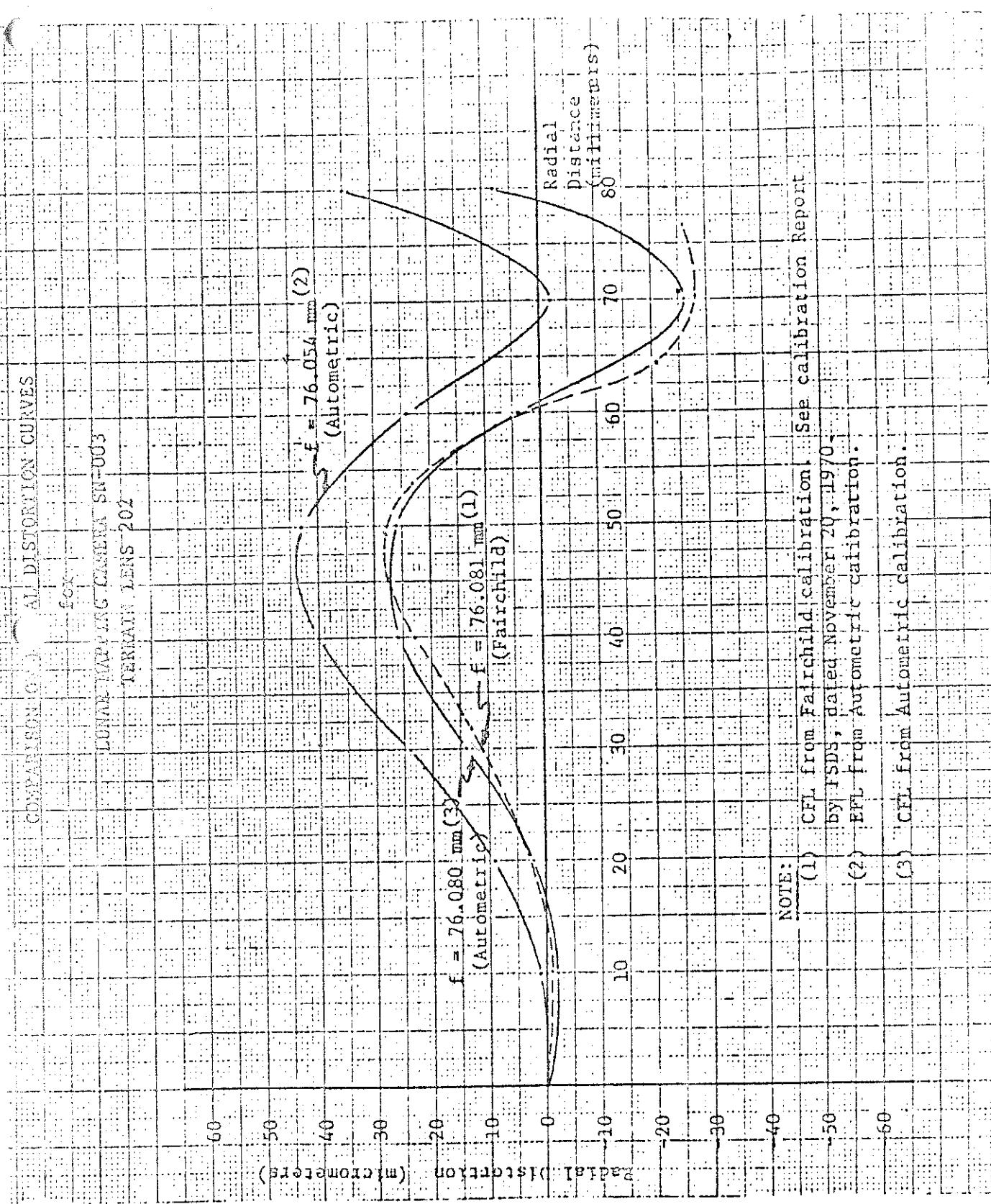


Figure 3.

3. Relative Orientation System:

The relative orientation matrix

$$M = \begin{bmatrix} m_{11} & m_{12} & m_{13} \\ m_{21} & m_{22} & m_{23} \\ m_{31} & m_{32} & m_{33} \end{bmatrix}$$

gives the angular orientation of the stellar camera coordinate system with respect to the terrain camera coordinate system. The orientation matrix can be factored into three orthogonal matrices each representing a simple rotation of the stellar camera coordinate system about a particular stellar axis. The sequence of the three rotations must be specified, because different angular orientations result from different sequences. The orientation of  $x_s$ ,  $y_s$ ,  $z_s$  with respect to  $X_T$ ,  $Y_T$ ,  $Z_T$  can be developed as follows.

Consider a stellar camera coordinate system  $x$ ,  $y$ ,  $z$  initially coincident with the terrain camera coordinate system  $X_T$ ,  $Y_T$ ,  $Z_T$  (refer to Figure 4). The three rotations  $\omega$ ,  $\phi$ ,  $\kappa$  are applied to the stellar camera coordinate axes in the given sequence to place the system into its final position,  $x_s$ ,  $y_s$ ,  $z_s$ .

- $\omega$  (roll) - Rotation about the  $x$  axis. Positive  $\omega$  takes the  $+y$  axis toward the  $+z$  axis, resulting in  $x'$ ,  $y'$ ,  $z'$  in Figure 4.
- $\phi$  (pitch) - Rotation about the  $y'$  axis. Positive  $\phi$  takes the  $+z'$  axis toward the  $+x'$  axis, resulting in  $x''$ ,  $y''$ ,  $z''$  in Figure 4.
- $\kappa$  (yaw) - Rotation about the  $z''$  axis. Positive  $\kappa$  takes the  $+x''$  axis toward the  $+y''$  axis, resulting in the final position of the stellar camera coordinate system  $x_s$ ,  $y_s$ ,  $z_s$  in Figure 4.

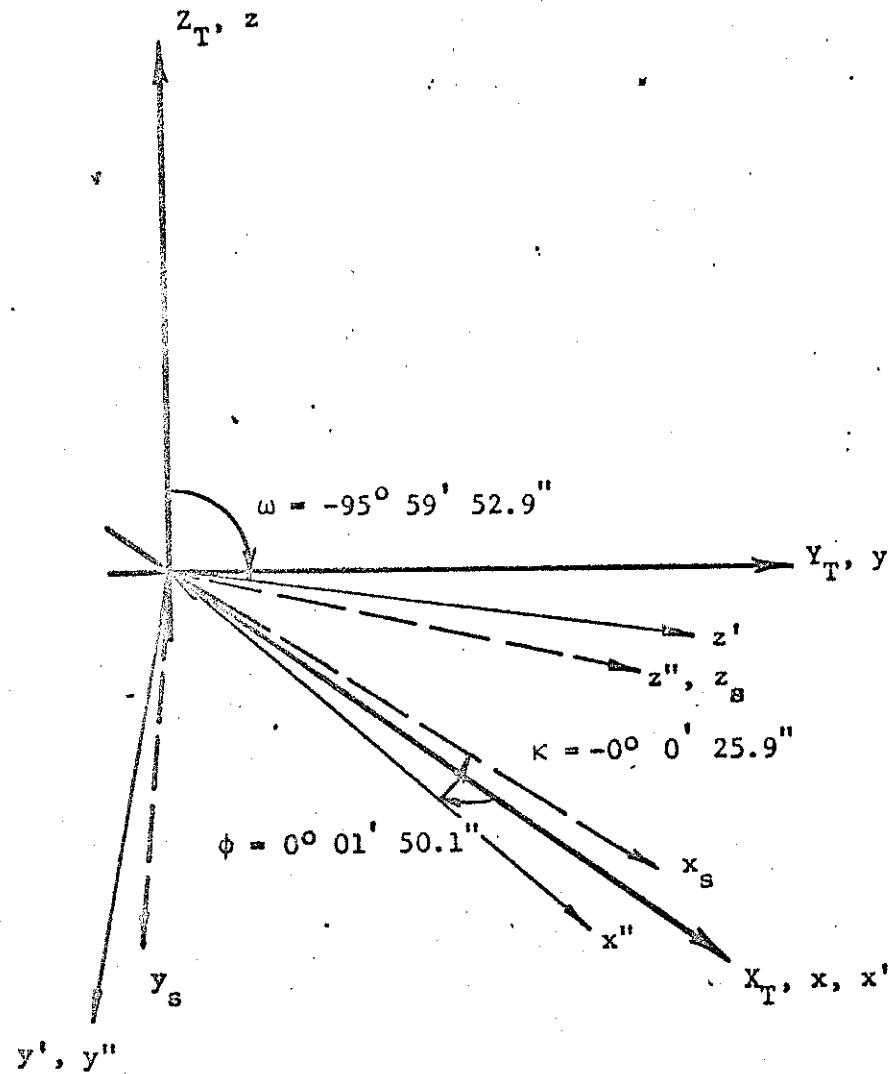


Figure 4. Orientation of Stellar Camera Coordinate System  
With Respect to Terrain Camera Coordinate System.

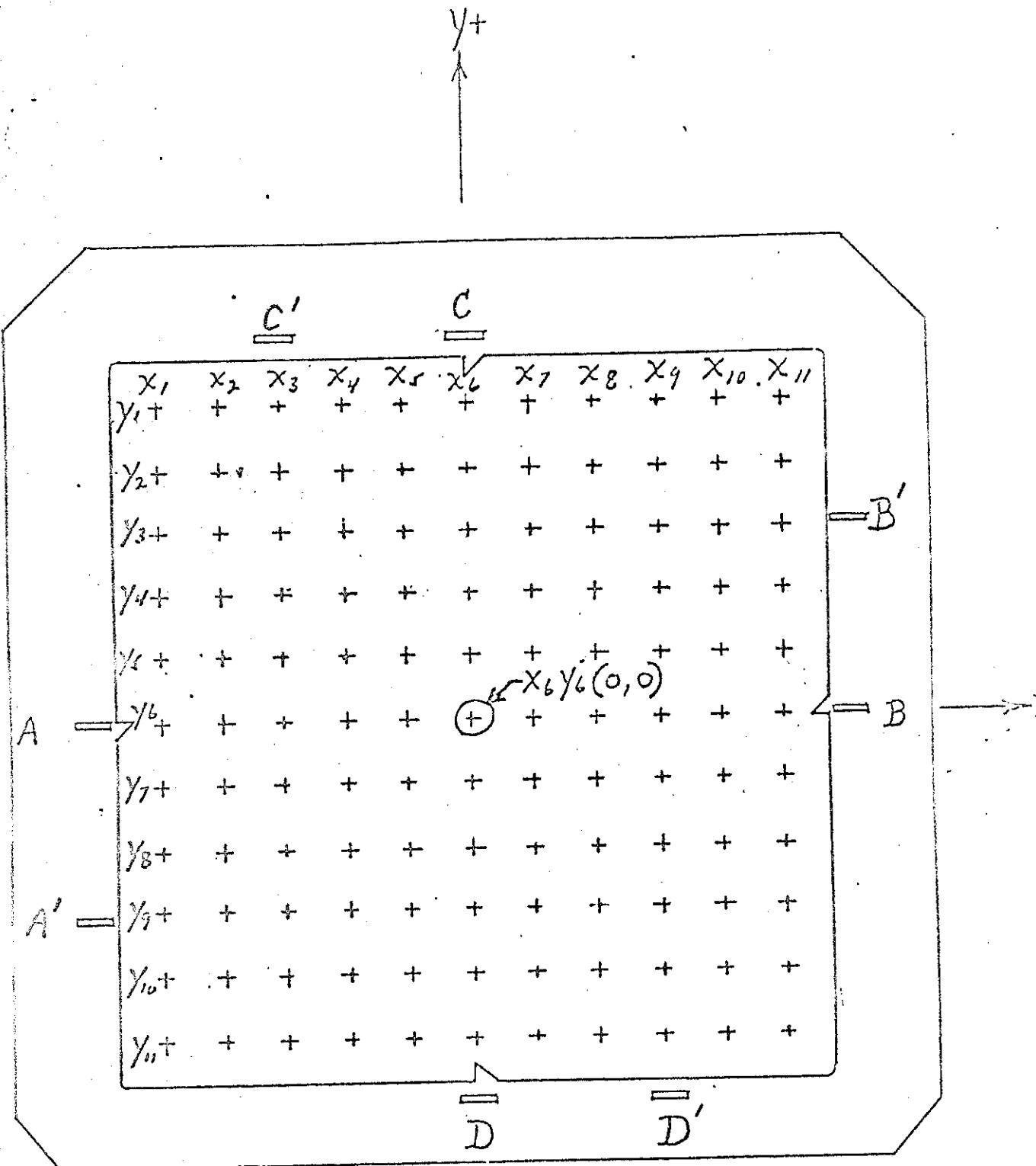
This Final Report was prepared for Fairchild Space and Defense Systems by Raytheon Company, Autometric Operation, under Contract N-0234.

*Ronald K. Brewer*

Ronald K. Brewer  
Senior Scientist/Photogrammetry  
Program Manager

STELLAR CALIBRATION REPORT SUPPLEMENT

Calibrated Ikogon (Mapping)  
and  
Ikotar (Stellar)  
Reseau Coordinate Intersections



{EMULSION UP}

$11 \times 11$   
reseaux matrix

LINE OF FLIGHT

negative?

IKOGON B MASTER GRID PLATE  
CALIBRATION DATA  
( COORDINATES IN MILLIMETERS )

CROSS	COORDINATES		CROSS	COORDINATES	
	X-	Y-		X-	Y-
X1, Y1	-50.0014	+50.0019	X1, Y4	-50.0014	+20.0015
X2, Y1	-39.9997	+50.0019	X2, Y4	-39.9997	+20.0015
X3, Y1	-29.9977	+50.0019	X3, Y4	-29.9977	+20.0015
X4, Y1	-19.9974	+50.0019	X4, Y4	-19.9974	+20.0015
X5, Y1	- 9.9990	+50.0019	X5, Y4	- 9.9990	+20.0015
X6, Y1	0.0000	+50.0019	X6, Y4	0.0000	+20.0015
X7, Y1	+ 9.9992	+50.0019	X7, Y4	+ 9.9992	+20.0015
X8, Y1	+19.9995	+50.0019	X8, Y4	+19.9995	+20.0015
X9, Y1	+29.9994	+50.0019	X9, Y4	+29.9994	+20.0015
X10, Y1	+40.0019	+50.0019	X10, Y4	+40.0019	+20.0015
X11, Y1	+50.0029	+50.0019	X11, Y4	+50.0029	+20.0015
X1, Y2	-50.0014	+40.0022	X1, Y5	-50.0014	+10.0013
X2, Y2	-39.9997	+40.0022	X2, Y5	-39.9997	+10.0013
X3, Y2	-29.9977	+40.0022	X3, Y5	-29.9977	+10.0013
X4, Y2	-19.9974	+40.0022	X4, Y5	-19.9974	+10.0013
X5, Y2	- 9.9990	+40.0022	X5, Y5	- 9.9990	+10.0013
X6, Y2	0.0000	+40.0022	X6, Y5	0.0000	+10.0013
X7, Y2	+ 9.9992	+40.0022	X7, Y5	+ 9.9992	+10.0013
X8, Y2	+19.9995	+40.0022	X8, Y5	+19.9995	+10.0013
X9, Y2	+29.9994	+40.0022	X9, Y5	+29.9994	+10.0013
X10, Y2	+40.0019	+40.0022	X10, Y5	+40.0019	+10.0013
X11, Y2	+50.0029	+40.0022	X11, Y5	+50.0029	+10.0013
X1, Y3	-50.0014	+30.0030	X1, Y6	-50.0014	0.0000
X2, Y3	-39.9997	+30.0030	X2, Y6	-39.9997	0.0000
X3, Y3	-29.9977	+30.0030	X3, Y6	-29.9977	0.0000
X4, Y3	-19.9974	+30.0030	X4, Y6	-19.9974	0.0000
X5, Y3	- 9.9990	+30.0030	X5, Y6	- 9.9990	0.0000
X6, Y3	0.0000	+30.0030	X6, Y6	0.0000	0.0000
X7, Y3	+ 9.9992	+30.0030	X7, Y6	+ 9.9992	0.0000
X8, Y3	+19.9995	+30.0030	X8, Y6	+19.9995	0.0000
X9, Y3	+29.9994	+30.0030	X9, Y6	+29.9994	0.0000
X10, Y3	+40.0019	+30.0030	X10, Y6	+40.0019	0.0000
X11, Y3	+50.0029	+30.0030	X11, Y6	+50.0029	0.0000

IKOGON B MASTER GRID PLATE  
CALIBRATION DATA ----- (COORDINATES IN MILLIMETERS)

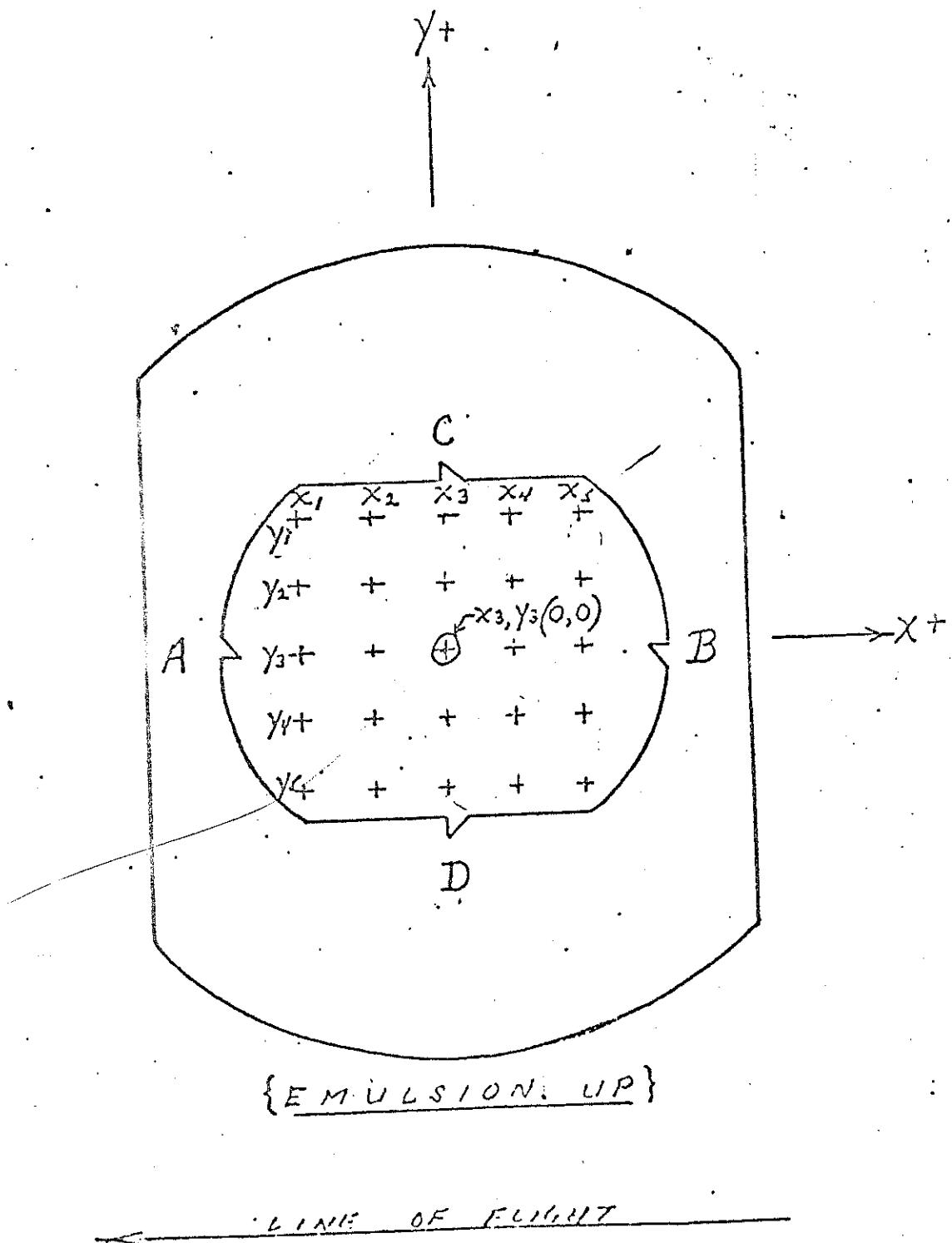
CROSS	COORDINATES		CROSS	COORDINATES	
	X-	Y-		X-	Y-
X1, Y7	-50.0014	-10.0000	X1, Y10	-50.0014	-39.9974
X2, Y7	-39.9997	-10.0000	X2, Y10	-39.9997	-39.9974
X3, Y7	-29.9977	-10.0000	X3, Y10	-29.9977	-39.9974
X4, Y7	-19.9974	-10.0000	X4, Y10	-19.9974	-39.9974
X5, Y7	- 9.9990	-10.0000	X5, Y10	- 9.9990	-39.9974
X6, Y7	0.0000	-10.0000	X6, Y10	0.0000	-39.9974
X7, Y7	+ 9.9992	-10.0000	X7, Y10	+ 9.9992	-39.9974
X8, Y7	+19.9995	-10.0000	X8, Y10	+19.9995	-39.9974
X9, Y7	+29.9994	-10.0000	X9, Y10	+29.9994	-39.9974
X10, Y7	+40.0019	-10.0000	X10, Y10	+40.0019	-39.9974
X11, Y7	+50.0029	-10.0000	X11, Y10	+50.0029	-39.9974
X1, Y8	-50.0014	-19.9992	X1, Y11	-50.0014	-49.9969
X2, Y8	-39.9997	-19.9992	X2, Y11	-39.9997	-49.9969
X3, Y8	-29.9977	-19.9992	X3, Y11	-29.9977	-49.9969
X4, Y8	-19.9974	-19.9992	X4, Y11	-19.9974	-49.9969
X5, Y8	- 9.9990	-19.9992	X5, Y11	- 9.9990	-49.9969
X6, Y8	0.0000	-19.9992	X6, Y11	0.0000	-49.9969
X7, Y8	+ 9.9992	-19.9992	X7, Y11	+ 9.9992	-49.9969
X8, Y8	+19.9995	-19.9992	X8, Y11	+19.9995	-49.9969
X9, Y8	+29.9994	-19.9992	X9, Y11	+29.9994	-49.9969
X10, Y8	+40.0019	-19.9992	X10, Y11	+40.0019	-49.9969
X11, Y8	+50.0029	-19.9992	X11, Y11	+50.0029	-49.9969
X1, Y9	-50.0014	-29.9987			
X2, Y9	-39.9997	-29.9987			
X3, Y9	-29.9977	-29.9987			
X4, Y9	-19.9974	-29.9987			
X5, Y9	- 9.9990	-29.9987			
X6, Y9	0.0000	-29.9987			
X7, Y9	+ 9.9992	-29.9987			
X8, Y9	+19.9995	-29.9987			
X9, Y9	+29.9994	-29.9987			
X10, Y9	+40.0019	-29.9987			
X11, Y9	+50.0029	-29.9987			

CERTIFICATION

*J. C. M.*

Signed

# EKOTAR B GRID PLATE LAYOUT



IKOTAR B MASTER GRID PLATE  
CALIBRATION DATA  
(COORDINATES IN MILLIMETERS)

CROSS	COORDINATES		CROSS	COORDINATES	
	X-	Y-		X-	Y-
X1, Y1	-10.0015	+10.0003	X1, Y4	-10.0015	-5.0002
X2, Y1	-5.0004	+10.0005	X2, Y4	-5.0008	-5.0005
X3, Y1	-0.0003	+10.0000	X3, Y4	-0.0005	-5.0004
X4, Y1	+4.9996	+10.0000	X4, Y4	+4.9995	-5.0000
X5, Y1	+9.9996	+10.0000	X5, Y4	+9.9996	-5.0002
X1, Y2	-10.0015	+5.0011	X1, Y5	-10.0015	-10.0015
X2, Y2	-5.0004	+5.0001	X2, Y5	-5.0010	-10.0015
X3, Y2	0.0000	+5.0000	X3, Y5	-0.0005	-10.0015
X4, Y2	+4.9996	+5.0006	X4, Y5	+4.9992	-10.0015
X5, Y2	+9.9992	+5.0004	X5, Y5	+9.9992	-10.0015
X1, Y3	-10.0015	+0.0006			
X2, Y3	-5.0009	+0.0002			
X3, Y3	0.0000	+0.0000			
X4, Y3	+4.9993	+0.0006			
X5, Y3	+9.9992	+0.0004			

~~FAIRCHILD~~ ~~SUPER 8MM FILM PROCESSOR~~  
A division of Fairchild Camera and Instrument Corporation  
El Segundo, California

LENS TYPE IKOTAR "B"  
SERIAL NUMBER 104

DATE 6 Oct. 1970

11. RESOLVING POWER

FILM 8400

PROCESSING Mx641 1 min @ 69° DIAG. #1 AWAR 84.4

TARGET CONTRAST 1000:1 DIAG. #2 AWAR 84.6

DIRG. #1	0°	2 $\frac{1}{2}$ °	5°	7 $\frac{1}{2}$ °	10°							
RADIAL	90	90	69	84	65							
TANGENTIAL	90	85	89	79	75							

DIRG. #2	0°	2 $\frac{1}{2}$ °	5°	7 $\frac{1}{2}$ °	10°							
RADIAL	90	75	95	84	75							
TANGENTIAL	90	95	85	79	78							

FAIRCHILD SPACE AND DEFENSE SYSTEMS  
A DIVISION OF FAIRCHILD CAMERA AND INSTRUMENT CORPORATION

Lens Type MKOGON "B"

Date 12 Oct 70

Serial Number 202

Relative Illumination

Axis	<u>5°</u>	<u>10°</u>	<u>15°</u>	<u>20°</u>	<u>25°</u>	<u>30°</u>	<u>35°</u>	<u>40°</u>	<u>45°</u>
98%	98%	100%	99%	98%	96%	91%	78%	66%	45%

$$\sum = 84.9\%$$

Lens Transmittance = 43.9%

$$\text{Lens T/F} = \frac{4.5}{\sqrt{.439}} = \frac{4.5}{.663} = 6.787$$

$$\text{Apert.} = \frac{6.787}{\sqrt{.049}} = \frac{6.787}{.922}$$

$$\text{Apert.} = 7.36$$

Relative Aperture

$$\frac{\text{FL}}{\text{EFFECTIVE APERTURE}} = \frac{2.0035''}{.690''} = 4.35$$

EXPOSURE Date: Nov 20, 1970

3.1 RESOLVING POWER - CONTRAST 1000:1

3601 3601

17.5°	22.5°	27.5°	30°	32.5°	35°	37.5°	40°	42.5°
123	120	124	105	126	134	134	129	106
126	123	124	162	162	135	162	158	130

3.3 135.5

L/MM

A DIVISION OF FAIRCHILD CAMERA & INSTRUMENT CORPORATION

EL SEGUNDO, CALIFORNIA

Lens Type IKOTAR "B"

Date 6 Oct. 1970

Serial Number 104

Relative Illumination

<u>10°</u>	<u>20°</u>	<u>30°</u>	<u>40°</u>	<u>Axis</u>	<u>220</u>	<u>50°</u>	<u>70°</u>	<u>10°</u>
90	96	99	100	100	100	97	94	88

Transmission

93%

Relative Aperture

f/2.75

## 6.2 CALIBRATION DATA FOR THE OPTICAL BAR PANORAMIC CAMERA

The ITEK calibration reports include relative illumination, spectral transmittance, veiling glare, calibrated focal length and system resolution.

# TEST PROCEDURE

FOR  
PROJECT 9446  
RELATIVE ILLUMINATION  
MEASUREMENTS  
PANORAMIC CAMERA  
FOR SCIENTIFIC INSTRUMENT  
MODULE

EXPERIMENT S-163

Itek

ITEK CORPORATION

Lexington 73, Massachusetts

Date 9-15-70

Long SN N-54

	PREPARED	PROJECT APPROVAL	QUALITY ASSURANCE APPROVAL
By	R. SHERLOCK	C. BACKE	R. WESPISER
Signed	<i>R. Sherlock</i>	<i>C. Backe</i>	<i>R. Wespiser SA</i>
Date	9/17/70	9/18/70	9/18/70

CUST./GOV'T. REP. \_\_\_\_\_ Date \_\_\_\_\_  
Reviewed \_\_\_\_\_

## S.3.1.3 Itek Test Data Sheet

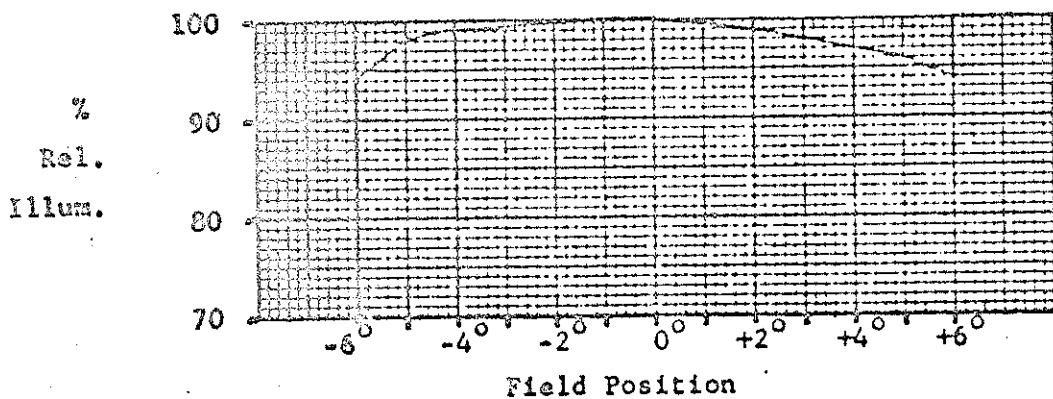
## RELATIVE ILLUMINATION MEASUREMENT

Panoramic Camera Lens, P/N 105150, Serial No. N-54

Field Position	Radiometer Readings vdc	Relative Illumination %
<b>B</b>	-6°	76.0
	-5°	77.5
	-4°	80.0
	-3°	80.5
	-2°	80.2
	-1°	81.0
<b>A</b>	0°	81.0
<b>B</b>	+1°	80.5
	+2°	80.0
	+3°	79.5
	+4°	78.5
	+5°	77.5
	+6°	76.0

Meter Scale for Radiometer Readings: 100-

$$\% \text{ Relative Illumination} = \frac{B}{A} \times 100$$



Data Recorded By: Itek/JS/6 Date: 10/12/70  
 QA Monitor: A.C. Kelley 10/15/70  
 Project Approval: Carl Baile 10/19/70

Test Procedure No. TP125

Q199-5 11/65

Page 3ITEK CORPORATION  
Lexington 73, Massachusetts

**TEST PROCEDURE**

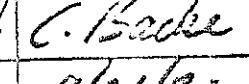
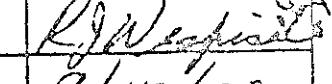
FOR  
PROJECT 9446  
SPECTRAL TRANSMITTANCE  
MEASUREMENTS  
PANORAMIC CAMERA  
FOR SCIENTIFIC INSTRUMENT  
MODULE  
EXPERIMENT S-163

The logo consists of the word "Itek" in a bold, sans-serif font, enclosed within a rounded rectangular border.**ITEK CORPORATION**

Lexington 73, Massachusetts

Date 9-15-70

Lem S/N N-84

	PREPARED	PROJECT APPROVAL	QUALITY ASSURANCE APPROVAL
By	R. SHERLOCK	C. BACKE	R. WESPISER
Signed			
Date	9/17/70	9/21/70	9/17/70

CUST./GOV'T. REP.

Date

Reviewed

## 6.3.1.3 Itek Test Data Sheet

## SPECTRAL TRANSMITTANCE MEASUREMENT

Panoramic Camera Lens, P/N 105150, Serial No. N-54

Wavelength nm	Radiometer Readings				% Transmittance
	A	A <sub>1</sub>	B	B <sub>1</sub>	
400	9.2	2.65	2.0	.13	22.6
420	25.8	6.6	5.9	.55	39.8
440	46.0	12.2	9.5	1.3	51.6
480	100.0	27.0	21.0	3.8	67.0
520	144.4	41.0	32.5	6.6	71.8
560	180.0	44.5	36.0	7.4	73.9
601	145.0	39.5	33.0	6.2	69.0
640	125.0	33.0	21.5	4.8	63.8
680	106.0	25.7	19.0	3.5	56.7
720	80.0	18.0	15.0	2.0	49.4

## Legend:

A = Brightness of the calibrated standard Lambertian source using the radiometer telescope.

A<sub>1</sub> = Brightness of the collimator target as seen from the lens test position using the radiometer telescope.

B = Brightness of the calibrated standard Lambertian source using the radiometer microscope.

B<sub>1</sub> = Brightness of the collimator target aerial image at the image plane (lens in place) using the radiometer microscope.

$$\% \text{ Transmittance} = \left[ \frac{B_1}{B} \div \frac{A_1}{A} \right] \times 100.$$

Data Recorded By: 6/12/70 Date: 6/12/70QA Monitor: G.C. W.H. May Date: 6-15-70Project Approval: John S. Miller Date: 6/18/70Test Procedure No. TP123Page 8

Q199-5 11/65

ITEK CORPORATION  
Lexington 73, Massachusetts

## 6.3.1.3 Itek Test Data Sheet

## T STOP CALCULATION

Panoramic Camera Lens, P/N 105150, Serial No. N-54

Wavelength nm	% Transmittance	T STOP
400	22.6	7.4
420	39.8	5.5
440	51.6	4.9
480	67.0	4.3
520	71.8	4.1
560	73.7	4.1
601	69.0	4.2
640	63.8	4.4
680	56.7	4.6
720	49.4	5.0

## Legend:

$$T \text{ STOP} = \frac{f/\text{number}}{\sqrt{t}} \quad \text{where } f/\text{number} \text{ is 3.5 and}$$

t is transmittance.

Data Recorded By: Eugen J. StollDate: OCT 12 1970QA Monitor: G.C. DeKey10-15-70Project Approval: Carl Basile10-19-70Test Procedure No. TP123Page 9

Q199-5 11/65

ITEK CORPORATION  
Lexington 73, Massachusetts

**TEST PROCEDURE**

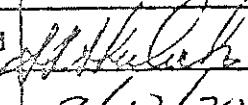
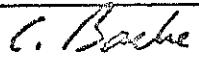
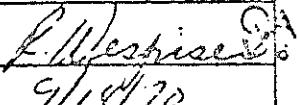
**FOR**  
PROJECT 9446  
VEILING GLARE MEASUREMENTS  
PANORAMIC CAMERA  
FOR SCIENTIFIC INSTRUMENT  
MODULE  
EXPERIMENT S-163

The logo consists of the word "Itek" in a bold, lowercase, sans-serif font, enclosed within a thin rectangular border.

ITEK CORPORATION  
Lexington 73, Massachusetts

Date 9-15-70

Item S/N N-54

	PREPARED	PROJECT APPROVAL	QUALITY ASSURANCE APPROVAL
By	R. SHERLOCK	C. BACKE	R. WESPISER
Signed			
Date	9/17/70	9/21/70	9/18/70

CUST./GOV'T. REP. \_\_\_\_\_ Date \_\_\_\_\_

Reviewed \_\_\_\_\_

## 6.3.1.3 Itek Test Data Sheet

## VEILING CLARE MEASUREMENT

Panoramic Camera Lens, P/N 105150, Serial No. N-54

	Step Wedge							Black Dot	% Veiling Glare
	Step No.	1	2	3	4	5	6	7	
* Calibrated Value									
		1.22	1.07	.92	.78	.65	.50	.36	
* Measured Test Values									
	-6°	.24	.34	.51	.30	1.14	1.44	1.77	.55
	-4°	.22	.30	.43	.65	.94	1.26	1.62	.44
	-2°	.60	.93	1.30	1.62	1.88	2.04	2.24	1.34
	0°	.24	.32	.50	.80	1.13	1.44	1.74	.54
	+2°	.26	.36	.55	.84	1.11	1.40	1.77	.82
	+4°	.25	.37	.56	.86	1.14	1.45	1.83	.84
	+6°	.22	.30	.46	.71	1.05	1.36	1.70	.64

Legend: \* Density values are logarithmic as read on the Macbeth Densitometer.

% Veiling Glare is computed from the black spot density plot (attached) for each field position.

Data Recorded By: Itek/Italk Date: 10/12/70

QA Monitor: JK/B/SB 10/16/70

Project Approval: Couf/Beck 10/19/70

Test Procedure No. TP127

No. of Pages 5

TEST PROCEDURE  
FOR  
PROJECT 9446  
C.F.L./MEASUREMENTS  
PANORAMIC CAMERA  
FOR SCIENTIFIC INSTRUMENT  
MODULE  
EXPERIMENT S-163

Itek

ITEK CORPORATION  
Lexington 73, Massachusetts

Date 9-15-70

PREPARED		PROJECT APPROVAL	QUALITY ASSURANCE APPROVAL
By	Signed	C. BACKE	R. WESPISER
Date	<u>9/17/70</u>	<u>9/21/70</u>	<u>9/21/70</u>

CUST./GOV'T. REP. \_\_\_\_\_ Date \_\_\_\_\_

Reviewed

Q199-1 11/65

3

## 6.3.1.3 Test Data Summary

Lens - N-54

## C.F.L. Calibration Summary

Filter	Field Position (degrees)	Mean C.F.L. (inches)	STD Deviation of of mean (C.F.L.)
23A	+6	24.0057	0.0004
	+4	24.0058	0.0003
	+2	24.0015	0.0006
	0 *	24.0052	0.0008
	-2	24.0073	0.0006
	-4	24.0057	0.0001
	-6	24.0050	0.0003
12	+6	24.0063	0.0008
	+4	24.0080	0.0003
	+2	24.0086	0.0009
	0 *	24.0081	0.0004
	-2	24.0088	0.0006
	-4	24.0089	0.0004
	-6	24.0080	0.0001
8	+6	24.0076	0.0004
	+4	24.0096	0.0007
	+2	24.0106	0.0005
	0 *	24.0080	0.0007
	-2	24.0059	0.0003
	-4	24.0081	0.0015
	-6	24.0064	0.0009
2A	+6	24.0044	0.0003
	+4	24.0075	0.0006
	+2	24.0085	0.0008
	0 *	24.0062	0.0007
	-2	24.0047	0.0013
	-4	24.0055	0.0006
	-6	24.0066	0.0005
no filter	+6	24.0037	0.0003
	+4	24.0050	0.0004
	+2	24.0065	0.0006
	0 *	24.0045	0.0005
	-2	24.0032	0.0004
	-4	24.0048	0.0008
	-6	24.0037	0.0006

\* Average of all field positions

Quality Assurance Review

A. C. McKey

QA

F32

31003

1-27-71

# ITEK APPROVED

## TEST PROCEDURE

### FOR

PHOTOGRAPHIC BASELINE (PBL)

(DYNAMIC TESTING WITH FORWARD MOTION COMPENSATION)

ACCEPTANCE OF THE PANORAMIC CAMERA

EXPERIMENT S-163

LOSS 23A Filter.

ITEK CORPORATION

Lexington 73, Massachusetts

Date 9/23/70

	PREPARED	PROJECT APPROVAL	QUALITY ASSURANCE APPROVAL
By	D.D. Auer	J. R. H.	W. ZEBLEY
Signed	A. E. Taube	J. F. L.	W. Zebley
Date	7/24/70	7/24/70	7/24/70

CUST./GOV'T. REP.

Date

Reviewed

Q199-1 11/63

No FILTER  
FLIGHT CONFIG.

79

DYNAMIC TESTING WITH MIC 10% SCALE  
DATA SHEET

FORWARD POSITION 0°

EPO (46)			EPO (47)			EPO (48)			EPO (49)			EPO (50)			EPO (51)			EPO (52)			
No.	Collimator Res. Ang.	Reed Lattice	COLLIMATOR			COLLIMATOR			COLLIMATOR			COLLIMATOR			COLLIMATOR			COLLIMATOR			
			Res.	Ang.	Res.	Res.	Ang.	Res.	Res.	Ang.	Res.	Res.	Ang.	Res.	Res.	Ang.	Res.	Res.	Ang.		
1	111/111	111/111	108/120	102/110	136/121	59	141/123	126	152/152	52	72/76	73	81/69	73	51	76	64	111/111	111/111	111/111	
2	136/104	121	128/117	120	128/128	152	152/152	152	121/151	37	91	156	44	59/81	58	72	74	57	55	77	
3	191/161	153	148/122	136	121/150	32	161/152	151	132	44	106/111	120	76/91	83	68/76	72	62	62	68	63	
4	152/149	142	149/151	163	119/125	121	152/111	162	129/152	52	150/119	121	113/76	76	68/72	70	53	53	53	53	
5	123/134	136	152/131	157	111/117	111	152/123	111	152/131	111	150/108	131	135/76	60/92	60/62	27	45	45	45	42	
6	136/121	125	121/121	121	128/135	122	123/123	111	129/152	52	130/114	122	132/55	60/81	71	72	61	69	69	61	73
7	151/118	12	152/135	105	105/132	121	152/152	111	152/152	111	152/152	111	153/111	121	113/121	121	68/160	68/160	68/160	68/160	68/160
8	129/121	133	131/117	156	113/152	150	152/122	172	135/111	111	110/108	121	115/91	119	107/51	51	54	54	54	55	55
9	114/121	21	181/171	171	121/135	147	161/153	194	171/153	194	158/144	123	111/111	108/127	76/91	91	53	54	54	54	50
10	154/136	150	161/161	113	108/135	121	126/152	171	181/152	171	152/171	111	176/108	111	132/81	81	52	52	52	52	57
	AVERAGES	37		3		50		56		47		125		81		81		81		81	65

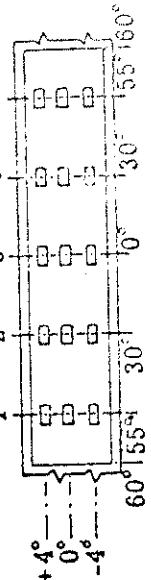
TP-70-9446-42-4

PAGE 38

*Blipotekan 6701*

*Telis J. G. Chutor*

Collimator positions



Nadir  
Format

DYNAMIC TESTING WITH THE VIBRO-SERIAL

DATA SHEET

RUN 951 G.

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## DYNAMIC TESTING WITH THE WLS RENE

## DATA SHEET

RUN 25147

Sister Branch of A

	COLLIMATOR 3 Ave.	COLLIMATOR 3 Resolution Ave.	COLLIMATOR 3 Ave.	COLLIMATOR 3 Resolution Ave.	COLLIMATOR 3 Ave.	COLLIMATOR 3 Resolution Ave.
PLATE NO.	III / $\equiv$	III / $\equiv$	III / $\equiv$	III / $\equiv$	III / $\equiv$	III / $\equiv$
1	120/100	120	121/102	104	122/102	102
2	131/103	135	144/132	131	132/132	130
3	144/132	149	154/132	149	152/132	145
4	144/132	129	141/129	127	123/125	126
5	144/132	129	141/132	129	132/132	129
6	121/126	129	122/124	129	121/122	126
7	120/126	126	121/124	123	121/122	125
8	136/125	127	148/127	125	132/126	127
9	161/152	157	125/132	125	132/132	127
10	128/132	129	132/132	132	121/121	129
AVERAGES						
	123	129	131	129	127	129

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Blepharobranch V3  
Left Mouth

Cut of fiber  
at very point  
6 point  
using - 40°

REQUEST NUMBER: RK1289

A C K N O W L E D G E M E N T S

When using the data in any reports, publications, or presentations,  
please acknowledge the National Space Science Data Center and the  
following individuals or groups:

71-063A-02D, 03D, 03L, 72-031A-02C, 03C, 03I, 72-096A-06B, 07C, 07J  
The Principal Investigator, Dr. Frederick J. Doyle.